

Review

# Chirality in Organic and Mineral Systems: A Review of Reactivity and Alteration Processes Relevant to Prebiotic Chemistry and Life Detection Missions

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**Abstract:** Chirality is a central feature in the evolution of biological systems, but the reason for biology's strong preference for specific chiralities of amino acids, sugars, and other molecules remains a controversial and unanswered question in origins of life research. Biological polymers tend toward homochiral systems, which favor the incorporation of a single enantiomer (molecules with a specific chiral configuration) over the other. There have been numerous investigations into the processes that preferentially enrich one enantiomer to understand the evolution of an early, racemic, prebiotic organic world. Chirality can also be a property of minerals; their interaction with chiral organics is important for assessing how post-depositional alteration processes could affect the stereochemical configuration of simple and complex organic molecules. In this paper, we review the properties of organic compounds and minerals as well as the physical, chemical, and geological processes that affect organic and mineral chirality during the preservation and detection of organic compounds. We provide perspectives and discussions on the reactions and analytical techniques that can be performed in the laboratory, and comment on the state of knowledge of flight-capable technologies in current and future planetary missions, with a focus on organics analysis and life detection.

**Keywords:** chirality; homochirality; origins of life; organics; minerals; biosignatures; life detection; prebiotic chemistry; enantiomeric excess



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## 1. Review Summary

This review article aims to bring together multiple perspectives on research centered around chiral determination in organic and mineral systems with a focus towards prebiotic chemistry and origins of life. Section 2 provides an introduction to the prebiotic chemistry field to give context and highlight the importance of understanding what chirality is and how it could have played a role in the origins of life. Included in this introduction is a brief summary and overview on the use of enantiomeric excess as a biosignature and the various hypotheses that have been proposed that led to the emergence of homochirality in biological systems.

The next part is separated into sections that review chirality in the two systems of focus: (1) organic systems, and (2) mineral systems. Sections 3 and 4 detail different types of organic chirality that can occur to highlight how chemical structure affects chiral properties and organic reactivity, especially in organic synthesis. This section draws heavily from examples in the pharmaceutical industry, which focus on organic synthesis/catalysis

research into asymmetric reactions. We present some examples of how chirality influences organometallic systems and their relevance to extraterrestrial materials and the origins of life. Here we present a table that provides examples of enantioselective reactions and asymmetric cross-coupling reactions relevant to various research disciplines. We summarize this section by highlighting the need to consider organic reactions within a geological framework, so that reaction products due to post-depositional geological alteration processes can be properly assessed and contextualized.

Section 5 discusses minerals that are relevant to prebiotic systems in the context of the composition of extraterrestrial bodies that may be explored during life detection missions and presents a table of the mineral and ice composition of major rocky planetary bodies in the Solar System. The chirality of mineral systems and its potential influence on organic systems and driving enantiomeric excess is discussed in Section 6. In particular, the section describes how mineral systems display a diverse range of chiral properties and introduces the concepts of enantiomorphic minerals, achiral minerals with chiral faces, local chiral sites, factors that influence a mineral's degree of enantioselectivity, mineral-organic interactions driving enantiomeric excess, as well as future directions for research in this field.

Section 7 of this review takes the geological context further by discussing geochemical processes involved in the preservation, alteration, and detection of organic matter as a primary and robust method for understanding the composition of primitive microorganisms as life began to evolve on Earth. In lieu of macroscopic, fossilized remains in the geological record, the study of early life involves examining molecular compositions, some of which have excellent preservation potential and can be traced back to specific organisms. Sections 7.1–7.3 detail the reactions and processes that occur during organic preservation and how chirality—particularly the stereochemical configuration of membrane lipids—changes during sequestration into refractory phases that get preserved over geologic time. In addition to terrestrial geochemical alteration processes, in Section 7.4 we discuss the parent body alteration processes that occur on asteroids and how the composition and chirality of organics are affected.

Section 8 summarizes analytical instrumentation and asymmetric solution phase and solid-state reactions. Section 8.1 includes a table that documents various separation and analytical methods for determining chirality and which of these has been proposed or flown on spaceflight missions. In the solution phase reactions, we detail three of the relevant reactions to prebiotic chemistry and origins of life: reductive amination (Section 8.2.1), Strecker synthesis (Section 8.2.2), and the formose reaction (Section 8.2.3). In Section 8.3 we focus on solid-state mechanochemical reactions, which have been recently shown to be viable options in the study of the origins of life. Additionally, we summarize some experimental studies on impact simulation that are directly related to the synthesis of amino acids, peptides, and molecules of prebiotic interest.

The final section of the review summarizes our recommendations for future research and discusses mission-relevant instrumentation for the analysis of organics and minerals and organic chirality (Section 9.1). We highlight the importance of continued instrument development (Section 9.2), contamination control (Section 9.3), and multidisciplinary collaboration (Section 9.4) in order to further the study of chirality in organic and mineral systems that is geared towards understanding the origins and contexts of how life evolved on Earth and where and how to find it elsewhere.

## 2. Prebiotic Chemistry, Chirality, and the Origins of Life

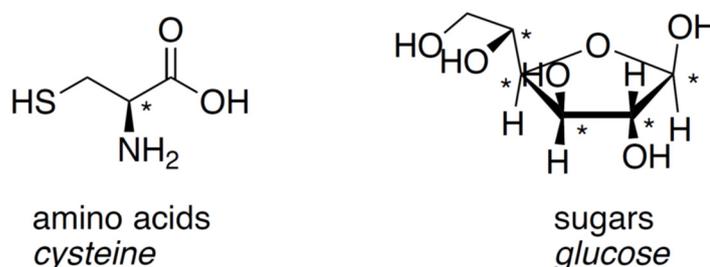
### 2.1. Chirality, Homochirality, and Enantiomeric Excess

Characterizing the organic chemical composition on planetary bodies is a key component in the search for habitable environments and evidence of extinct or extant life. The field of prebiotic chemistry focuses on how simple organic precursor compounds could have been synthesized and how they could have evolved into complex, functioning biomolecules, such as nucleic acids, peptides, and proteins [1–7]. These building block

precursors were likely required to carry out metabolic reactions and the information storage and transfer necessary for the maintenance and replication of living organisms [8,9].

Molecules can be arranged in three-dimensional space from a single point source (i.e., at a carbon atom). When a carbon atom has four distinct functional groups attached to it (i.e., is  $sp^3$  hybridized), the molecule can exist as one of two possible three-dimensional shapes, and that carbon atom is termed a “chiral center”. Molecular chirality—or handedness—whereby a molecule is not superimposable on its mirror image, is a property that is of great interest in the field of prebiotic chemistry. Chiral molecules are optically active; that is, they rotate plane-polarized light in opposite directions depending on their chirality. Enantiomers of chiral molecules possess identical chemical and physical properties, with the exception of how they interact with other chiral molecules or electromagnetic radiation. Thus, chirality can affect various properties of important molecules, such as amino acids and sugars (Figure 1), including: molecular self-assembly, asymmetric reactions/reactivity, molecular recognition and replication, and light, or spin polarization [10–13].

### Examples of chiral prebiotic molecules



**Figure 1.** Amino acids and sugars are examples of chiral molecules relevant to prebiotic chemistry. The asterisk (\*) indicates the chiral carbon center(s). (Left) is the amino acid cysteine ( $C_3H_7NO_2S$ ); (right) is the sugar glucose ( $C_6H_{12}O_6$ ).

Homochiral polymers are a result of the preference for one enantiomer over the other; biology on Earth has evolved to use homochiral proteins and nucleic acids, which are made up of L-amino acids and D-sugars, respectively [14,15], although there are rare exceptions [16]. Notably, life generally relies on L-amino acid and D-sugar monomers for metabolic functions as well. L- and D- are stereochemical designations for discerning enantiomeric pairs of amino acids and sugars, and are described in more detail in Section 3.1. Without a seed (i.e., a solid catalyst or already chiral reactant) that can induce asymmetry, chemical reactions will synthesize products in racemic mixtures, i.e., an equal mix of both enantiomers. Enantiomeric excess (ee) reflects the abundance of one particular enantiomer over the other and can be determined as a ratio of the observed specific rotation of the mixture over the specific rotation of the pure enantiomer, typically calculated as  $ee = [(L - D)/(L + D)] \times 100$  [17]. For example, an enantiopure product will have 100% of one enantiomer and 0% of the other, and racemic mixtures have ee of 0%, as there is no excess of one enantiomer over the other.

#### 2.2. Chirality in a Prebiotic Context

Before life arose on Earth, nucleosynthetic reactions occurring in the interstellar medium created elements that formed primitive organic compounds [18–22]. These organic compounds can be preserved in asteroids and comets and are delivered to Earth within carbonaceous chondrite meteorites [23]. Homochirality and the amplification of the enantiomeric imbalance resulting in the observed chirality of biological polymers has been hypothesized to have originated from either terrestrial or extraterrestrial processes before abiogenesis (i.e., the origins of life) or afterwards as a consequence of biological evolution [24]. Among the prebiotic hypotheses, there are theories that propose either

stand-alone or deterministic processes [15]. For example, the autocatalytic mathematical model developed by Frank [25] proposed that homochirality was amplified through the catalysis of a starting compound, which repeatedly synthesizes itself in one structural configuration while excluding the other [26]. Additionally, amplification in terrestrial environments of ee could have occurred with the aid of chiral reactants or substrates, as this has been demonstrated to promote the synthesis of specific enantiomers [27–29].

The effect of circularly polarized light and other energy sources, such as galactic cosmic rays (GCRs), on the chirality of organics is also well studied, e.g., Ref. [14] and references therein [30–32]. While prebiotic molecules of interest are generally synthesized in racemic mixtures, moderate abiotically synthesized ee have been measured in meteorites [14,17] and references therein. The organic composition of meteorites, specifically of hydroxy acids [33,34], typically demonstrate racemic composition and, in some cases of amino acids and sugar derivatives, meteoritic organics demonstrate an ee approaching that observed in biology [35–46].

### 2.3. Enantiomeric Excess as a Biosignature

Enantiomeric excess (ee) of organic molecules has been proposed as a biosignature based on our observations and knowledge of molecular biochemistry on Earth [17,47–49]. However, it remains to be observed whether ee is a robust indicator of life, given that it is unknown (1) if abiotic processes can generate significant enantiomeric excess such that it would be mistaken for a false positive for life detection; (2) whether the chemical evolution of near-homochiral polymers was required for life to evolve (if not, lack of ee could lead to false negatives); (3) if such chemistry is a requirement for all life or just Earth-like life; or (4) that biological systems always utilize a single enantiomer for building functional polymers. To expand on the fourth point, it remains possible for extraterrestrial life to rely on homochiral polymers of various chiralities; for example, perhaps it could use a protein composed of D-amino acids for one function and another composed of L-amino acids for another. Terrestrial life does this to some extent in that it incorporates only L-amino acids for building proteins, but uses D-amino acid monomers for various functions [50,51]; bacteria also use both D- and L-amino acids for the production of peptidoglycan [52], and some bacterial strains have even been demonstrated to be capable of growing using D-amino acids and L-sugars [53].

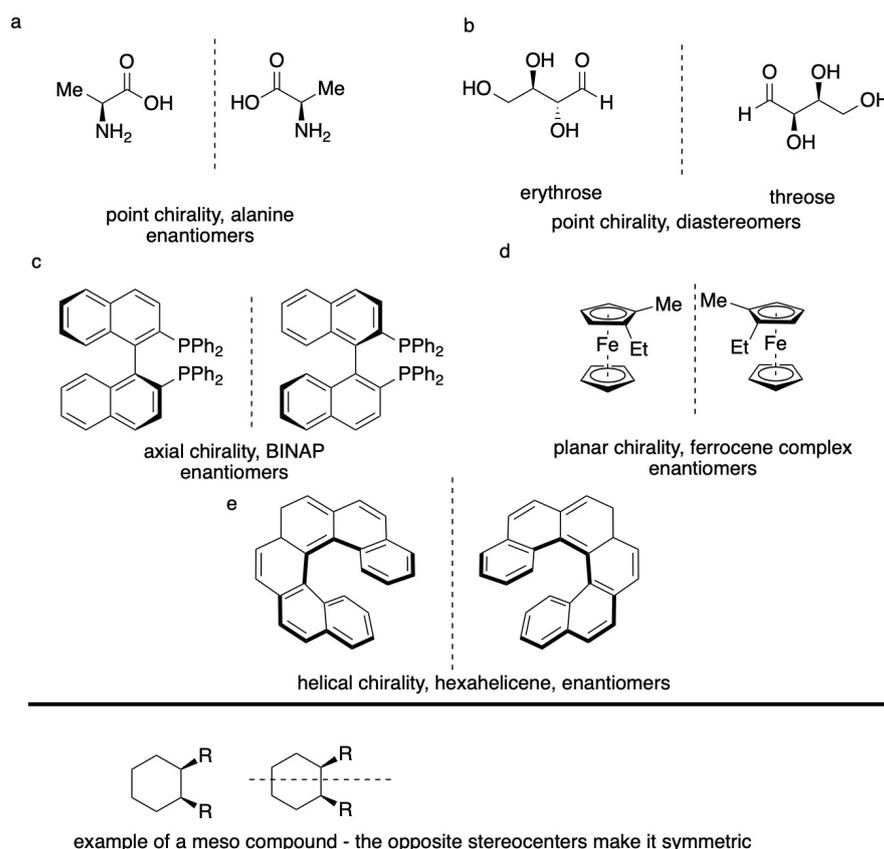
Understanding the pathways and environments relevant to chirality is critical to determining how homochiral polymers evolved on Earth and how we can utilize it as a possible biosignature to look for potential life on other planetary bodies. Specifically, the origins of life and astrobiology fields aim to constrain the geochemical conditions under which primitive/abiotic racemic or slightly enantioenriched mixtures of organic molecules could have chemically evolved towards homochiral/near homochiral polymers utilized in life as we know it.

In this paper, we provide a perspective on the interaction between chirality in organic systems with minerals that are relevant in prebiotic environments. We explore geochemical reactions and scenarios, such as the binding and release of organic compounds from mineral matrices during alteration and preservation, and the astrobiological implications of these phenomena. In the terrestrial context, alteration is defined as any reaction that occurs after the primary deposition of organic matter during sedimentary diagenesis and catagenesis. In the extraterrestrial context, alteration refers to processes occurring on the parent body. Examples include thermal or aqueous processes, shock wave propagation, and brecciation. We also discuss the current state of knowledge regarding the impact of minerals on driving different types of ee and provide some future experimental directions that could address gaps that warrant further study of organic and mineral chirality in the context of prebiotic chemistry and the origins of life.

### 3. Organic Chirality

#### 3.1. Point/Molecular Chirality

There are various types of chirality that organic molecules can possess; some different types are shown in Figure 2 (as illustrated by the amino acid alanine, Figure 2a). Point or molecular chirality is chirality that is based around an atom, such as carbon. Molecules that have atoms where each functional group attached to it is distinct are termed chiral [54]. For carbon, this typically means that the carbon atom has four different groups attached. Compounds that are mirror images of one another but cannot be superimposed through rotation are termed enantiomers; those that are non-superimposable, non-mirror images are termed diastereomers (example shown in Figure 2b).



**Figure 2.** Examples of point/molecular chirality in organic molecules. (a) Enantiomers of the amino acid alanine. (b) Diastereomers of erythrose (left) and threose (right). (c) Axial chirality exhibited by BINAP enantiomers. (d) Planar chirality of the enantiomers of a ferrocene complex. (e) Helical chirality of hexahelicene enantiomers. Bottom: generic dialkylcyclohexane as an example of a meso compound with an internal plane of symmetry, which is achiral and does not possess enantiomers.

Often, enantiomers are described by their absolute configuration as R (rectus, right) or S (sinister, left). The difference between R and S involves assigning priority to the different substituents and determining whether the identified lowest to highest priority groups follows a clockwise direction (designated R) or counterclockwise direction (designated S). This method is also called the Cahn–Ingold–Prelog rules in organic nomenclature [55]. The D/L designation is often used for sugars and amino acids; however, these do not directly line up with the R/S designation. In a Fischer projection of the molecule (for example, a sugar), the second to last carbon determines whether the molecule is D or L. If the hydroxyl group is on the left, the sugar is L, and if the hydroxyl group is on the right, the sugar is D. While the majority of chirality discussed in the context of origins of life is point/molecular chirality, there are other types of chirality that organics can possess, including axial, helical, and planar chirality, which we discuss below.

### 3.2. Heteroatom Chirality

This work will primarily focus on point/molecular chirality relevant to carbon; however, other atoms can be chiral, including P, N, and S, which are common heteroatoms in organic compounds. Additionally, other atoms within the carbon group, e.g., Ge and Si, can exhibit chiral structures [56–59]. In the case of heteroatoms, P, N, and S in their neutral form would typically need a lone pair of electrons and three different substituents attached to be considered chiral. A key for maintaining this configuration (i.e., chirality at the heteroatom) is to limit inversion around the lone pair. This is known as pyramidal inversion and, in the specific case of N, it is known as amine inversion. Amine inversion occurs rapidly at and below room temperature, and the N does not exhibit a chiral center. These are known as fluxional reactions and allow the enantiomers to racemize [60], which means pure enantiomers cannot be isolated. During amine inversion, in the transition state, there are three substituents arranged in the plane (co-planar) around the N atom, and the lone pair of electrons occupies an out-of-plane p orbital. Therefore, the structures containing chiral heteroatoms are often rigid or bulky, limiting the rate of inversion [61] and giving rise to enantiomers. Phosphine (PH<sub>3</sub>) and chiral sulfur-containing compounds, such as sulfites, sulfoxides, sulfonium salts, and sulfinic esters, also undergo the same inversion as amines, but the rate at room temperature is much lower [62,63]. Therefore, chiral P-phosphines and appropriately substituted chiral S compounds are optically stable at room temperature [64].

Compounds containing chiral heteroatoms can be used as ligands or reagents in enantioselective reactions and are also used in drugs [65]. Examples of chiral S-based drugs include armodafinil and esomeprazole, which, as sulfoxides, always have a stereogenic S center due to the lone pair of electrons [66,67]. These drugs typically have an enantiopure form and a racemic form that display distinct chemical properties and are marketed for different uses. There is also similar research being conducted on P chiral centers e.g., Ref. [68], as well as N [69,70]; however, the stereochemistry of the N atom makes it the most difficult heteroatom to control [70].

### 3.3. Additional Forms of Chirality

Molecules possessing features related to axial, helical [71], and planar [72] structure can also impart chirality in compounds that do not possess point chiral or stereogenic centers. Axial chirality is achieved when symmetry exists looking down an axis of a molecule; this commonly occurs in 2,2'-disubstituted biaryls and dissymmetrically substituted allenes. Examples of axial chirality include the organophosphorus compound, 2,2'-bis(diphenylphosphino)-1,1'-binaphthyl, or BINAP [73] and 2,2'-dihydroxy-4,4',6,6'-tetramethylbiphenyl. In both examples, axial chirality is due to restricted rotation as a result of steric hindrance (Figure 2c). Helical chirality is based on twisting in 3D space, as demonstrated in the aromatic compound hexahelicene (Figure 2d). Similar to axial chirality, steric strain promotes the existence of a stereogenic axis, giving rise to chiral forms in helically chiral structures. Planar chirality is a system where the 2D structures are not superimposable. They contain two dissymmetric, non-coplanar rings that cannot be easily rotated (e.g., metallocenes including ferrocene, Figure 2e).

In addition, not all structures with chiral components are chiral when the structure is examined as a whole. For example, if a meso compound has two or more chiral components (stereogenic centers), they “cancel out” the asymmetry, as the structure contains an internal plane of symmetry (Figure 2). An example of a meso compound is 2,3-dichlorobutane, which is not optically active nor does it have enantiomers. Achiral structures can also be prochiral, meaning that the addition of a unique functional group to a non-stereogenic center could give rise to chirality in an asymmetric reaction. For example, in an asymmetric Strecker reaction (see Table 1, entries 2 and 3), the imine group is prochiral. This means that depending on the face (i.e., Re or Si) that the reaction occurs on, different enantiomers (i.e., R and S) can be formed. In these mechanisms, other asymmetric components impart selectivity between the prochiral faces.

**Table 1.** Examples of enantioselective reactions and the research field(s) they are relevant to. The focus of this review is the reactions relevant to prebiotic chemistry and the origins of life.

Reaction	Starting Material	Product	Relevant Field(s)	References
Soai	pyrimidine-5-carbaldehyde	pyrimidyl alcohol	Origins of life; autocatalysis	[26,74]
Strecker	Aldehyde	Amino acid	Origins of life, pharmaceutical and natural product synthesis	[75]
Strecker	Ketone	$\alpha,\alpha$ -disubstituted amino acids	Origins of life, pharmaceutical and natural product synthesis	[76]
Reductive amination	Alpha keto acid	Amino acid	Origins of life, pharmaceutical and natural product synthesis	[77,78]
Kiliani–Fischer synthesis	Sugar	Monosaccharide	Origins of life	[79,80]
Sharpless epoxidation	Allylic alcohols	2,3-epoxyalcohols	Pharmaceutical and natural product synthesis	[81]
Sharpless bishydroxylation	Alkene	Vicinal diol	Pharmaceutical and natural product synthesis	[82]
Sharpless oxyamination	Alkene	Vicinal amino diol	Pharmaceutical and natural product synthesis	[83]
Midland reduction	Carbonyl (ketone)	Alcohol	Pharmaceutical and natural product synthesis	[84]
Noyori asymmetric hydrogenation	Keto ester	Hydroxy ester	Pharmaceutical and natural product synthesis	[85]
Corey-Itsuno reduction	Ketone (achiral)	Alcohol (chiral, non-racemic)	Pharmaceutical and natural product synthesis; industrial synthesis	[86,87]
Asymmetric Diels-Alder	Diene and alkene	Cyclohexene	Pharmaceutical and natural product synthesis; industrial synthesis	[88–93]
Examples of asymmetric cross-coupling reactions				
Suzuki-Miyaura	Alkyl- or arylhalides + organoborates	Alkyl or aryl compounds	Pharmaceutical and natural product synthesis; industrial synthesis; catalysis	[13,94,95]
Ni/Photoredox dual catalysis	Varied	Varied	Pharmaceutical and natural product synthesis; catalysis	[96–99]
Buckwald-Hartwig amination	Varied	Amine	Pharmaceutical and natural product synthesis; industrial synthesis; catalysis	[100,101]

### 3.4. Asymmetric Organic Synthesis Reactions

Asymmetric reactions relevant to the origins of life include reductive amination and Strecker synthesis for synthesizing amino acids and the autocatalytic formose reaction for synthesizing sugars (see Section 8.2.3 for more details). Asymmetric reactivity is very important in the field of organic chemistry, as it is often a key consideration in pharmaceutical and natural product synthesis [102,103]. Notably the 2021 Nobel Prize in Chemistry was awarded for the discovery of chiral organocatalysis [104]. There are a variety of catalytic methods to create asymmetry in reactions (Table 1). Some examples include organocatalysis that utilize chiral compounds (such as proline) as the catalyst, which then imparts asymmetry [105–107]. For metal-based catalysis, the chiral component can often be the ligand of the metal catalyst; such ligands can include organics with point or axial chirality [108–110]. However, there are other methods to impart chirality with the use of transition metals [111].

There are also chiral pool materials that are enantiopure compounds and readily and naturally available, such as amino acids and certain carboxylic acids, e.g., refs. [112,113]. Chiral pool materials are often commercially available and can be used in total synthesis. Total synthesis—common in organic chemistry—is the complete chemical synthesis of an organic molecule and often a complex natural product, such as taxol, strychnine, or a pharmaceutical. In these syntheses, organic chemists need to design enantioselective reactions to access complex chiral products.

In asymmetric organic synthesis reactions, geologic context is not often considered. Instead, considerations for chiral/asymmetric starting materials are often based on availability and yield of product. Ideally, scientists who utilize organic syntheses to impart asymmetry aim to synthesize their natural product with high purity and selectivity, therefore the conditions are often heavily optimized towards achieving those objectives. In the field of prebiotic chemistry, the geologic context is important, as yield and/or selectivity are not often the final goal of these projects, but instead to determine the conditions that drive product distribution [5,6].

#### 4. Organometallic Systems Related to Chirality

Organometallic compounds, whereby a carbon atom is covalently bonded to a metal, play important roles in enantioselective organic synthesis and methodology [114,115], industrial synthesis [116], materials science, nanotechnology [117,118], pharmaceuticals, and medicinal research [119,120], and in prebiotic and metabolically relevant systems [121–124] and references therein. Organometallic reagents can be used in other reactions as reagents e.g., R-MgBr (where R = alkyl, allyl, aryl, or vinyl group) for Grignard reactions [125,126] and catalysts, i.e., Pd and Ni compounds for Heck reactions [127–129]. Cross-coupling reactions are a notable type of organometallic reactions where two organic fragments are coupled together with the use of a metal catalyst, such as the Suzuki reaction [130,131]. Cross-coupling reactions and the use of precatalysts are powerful techniques that can be used to form C-C, C-N, and other C-X bonds, e.g., Refs. [132–134], to synthesize materials relevant to all fields of chemistry. While there are a significant number of different cross-coupling techniques, far more often than not, the techniques have not been explored for selectivity. Table 1 lists some of the examples where the reactions are enantioselective. In addition to interest in bioactive-starting materials relevant to the pharmaceutical industry, some cross-coupling reactions could have relevance in prebiotic environments. For example, the copper-mediated cross coupling of cyanide and acetylene synthesized the amino acids: arginine, aspartic acid, asparagine, aspartate, glutamine, glutamate, and proline [135].

Many proteins are associated with organometallic complexes or metalloenzyme cofactors [136], as they assist with the stabilization of their structure and support other vital biochemical reactions [123] and references therein. Organometallic compounds have been demonstrated to catalyze the polymerization of peptides and proteins [123,137,138]. The metal ions stabilize spectator ions during peptide formation within their ligands and the amino or carboxyl group of amino acids is protected during peptide synthesis [139,140].

Recently, there have been several reports of metal-organic compounds in meteoritic organic matter that have implications for understanding parent-body interactions with origins of life implications [141–143]. A study by Ruf and colleagues surveyed 61 meteorites over a wide range of petrologic types. They found dihydroxymagnesium carboxylates  $[(OH)_2MgO_2CR]^-$  in the soluble organic fraction that could be associated with metamorphic events, such as thermal alteration and shock events. These types of organometallic anion complexes had not been identified in meteorites prior to this study. Another study looked at the soluble organic fraction of 44 meteorites and found a novel homologous series of sulfur magnesium carboxylates. They appear to be thermally stable and their abundance was correlated with increasing thermal maturity [143]. Smith and colleagues looked at the origin of cyanide in the CM2 Murchison meteorite and found that it was primarily bound (and subsequently released) as iron cyanocarbonyl  $([Fe^I(CN)_5(CO)]^{3-})$

and  $[\text{Fe}^{\text{II}}(\text{CN})_4(\text{CO})_2]^{2-}$  organometallic complexes. These results suggest that cyanide in the form of iron cyanocarbonyl complexes could be a source for free cyanide delivered by meteorites and potentially be a precursor to catalytic centers of early enzymes, such as Fe hydrogenases [142]. Below, we describe mineral chirality and how organic reactions can impart chirality in organic-mineral systems in the context of prebiotically relevant environments.

## 5. Minerals Relevant to Prebiotic Chemistry

The composition of planetary surfaces varies based on the planet's histories, including its orbital position during early Solar System development, and subsequent physical, chemical, and space weathering processes (Table 2). 202 subdivided into broad categories of primary igneous minerals and secondary minerals produced by processes, including impacts, metamorphism, weathering, and volatile ices. Primary minerals include olivine, pyroxene, quartz, and feldspar. Secondary minerals include phyllosilicates, Fe oxides, carbonates, and sulfates. Common ices of volatile phases are driven by the "frost line" of that volatile phase during the early Solar System, or the distance at which that phase can condense into liquid form and resist further migration outward by the solar wind [143,144]. The abundance of water, ammonia, and methane ices are relatively low and dependent on regional conditions in the inner Solar System and more common from the Main Asteroid Belt outwards (Table 2).

**Table 2.** The surface minerals and ice surface compositions of major rocky bodies in the Solar System.

Planetary Body	Major Surface Minerals	Major Ices	References
Mercury	Plagioclase, olivine, pyroxene, sulfide, graphite	Water	[145–149]
Venus	Theorized: Olivine, pyroxene, sulfide, Fe oxides, carbonates, ilmenite, sulfate	None identified to date	[150,151]
Earth	Olivine, pyroxene, plagioclase, anorthite, quartz, Ca carbonate, phyllosilicates, Fe oxides	Water, lesser methane	[152–155]
Moon	Anorthite, plagioclase, pyroxene, olivine, ilmenite	Water	[146,156,157]
Mars	Olivine, pyroxene, phyllosilicates, sulfates, Fe oxides	Water, CO <sub>2</sub> , possibly methane	[146,158–161]
Asteroids, moons, and dwarf planets	Olivine, pyroxene, phyllosilicates, carbonates, Fe oxides	Water, methane, nitrogen, CO <sub>2</sub> , CO	[162–166]

Available surface minerals could have affected prebiotic chemistry in ways that are discernible from mineral records visible today. Because of their long-form structural repetition, available cations, and reactivity, many mineral structures have long been hypothesized to be important in prebiotic chemistry and possible emergence of life [167,168]. Indeed, mineral deposits control the availability of non-volatile bio-essential elements, so distinguishing planetary surface mineral deposits is critical to the question of habitability and preservation in the Solar System [168].

Clay minerals are a commonly cited class of minerals to be relevant to prebiotic chemistry, e.g., montmorillonite is found in carbonaceous chondrite meteorites [169,170]. Fe/Mg-clays are of particular interest when constraining planetary habitability, as they tend to be formed in alkaline, reducing environments that some studies propose are favorable for the transition from prebiotic to biotic activity [171–176], and may have even had a role in the origins of life on Earth [177]. The small particle size of clays could have concentrated and protected organic molecules from photolysis [178]. Studies have also suggested that amino acids could be polymerized in aqueous solution in the presence of clay minerals, e.g., montmorillonite [179]. Illite can promote chain elongation of amino

acids [180]. Hydrotalcite can serve to concentrate glycolaldehyde phosphate from dilute solution and thus catalyze condensation of the carbohydrate subunits [181].

Only some of these planetary minerals can structurally have different chiral forms. Mineral chirality is difficult to observe using the surface characterization methods traditional in planetary science, but as mineral chirality could have affected origins of life processes, a review of mineral chirality and how it intersects with prebiotic chemistry is of growing importance.

## 6. Mineral Chirality

Chirality is not just limited to organics; minerals can be intrinsically chiral (e.g., L versus D quartz), be overall achiral, but have chiral faces (e.g., calcite), have local chiral sites, or build chiral macromolecular structures (e.g., carbonate toroidal super structures [182] or spiraling of gastropod shells [183,184]). Mineral chirality is important to understand because interactions of organic molecules with mineral surfaces may also confer a degree of chiral selectivity in abiotic planetary systems.

### 6.1. Enantiomorphic Minerals

Mineral crystals are built from repeated translations of “unit cells,” or the smallest repeating pattern of atoms that reflects the symmetry and structure of the entire crystal. Unit cells are classified into one of six geometric crystal families, which groups crystals based on a combination of the unit cell’s shape (i.e., lattice structure) and the required symmetry of their point groups (i.e., a group that describes the symmetry operations under which the unit cell is invariant). Within each crystal family, the specific geometry and symmetry of a unit cell can be more specifically classified into space groups; thus, every mineral belongs to one of 230 space groups. The space group describes both the specific translational symmetry of the unit cell (i.e., the Bravais lattice, screw axis, and glide planes) and the point group symmetry operations (reflection, rotation, inversion, or rotoinversion). Minerals whose unit cells cannot be superimposed onto their mirror image are considered enantiomorphic (i.e., chiral); thus, asymmetry can be deduced based on the space group of the mineral. In other words, specific space groups are chiral and can be used to identify chiral minerals. Of the 230 possible space groups, 64 (belonging to 11 of the 32 crystal point groups) are chiral (Table 3). A full list of chiral minerals (as identified by their space group) and their chemical formula have been compiled (Table S1, Supplementary Material) using the mineral database Mindat. Notably, mineral enantiomorphs do not belong to the same space group.

As larger crystal structures reflect the inherent geometry and symmetry of the unit cell, minerals that are chiral at the unit cell will have crystal faces that are also chiral. Chiral minerals are composed of chiral unit cells which themselves may be made up of achiral subunits. For example, one of the most commonly discussed chiral minerals,  $\alpha$ -quartz, is composed of repeated molecules of achiral silicate ( $\text{SiO}_4$ ). It is the arrangement of the silicate molecules into left- or right-handed helices that give quartz its chirality (as evidenced by its unit cell and crystal lattice in Figure 3). Consequently, dissolution of chiral minerals will often result in achiral mixtures (e.g., dissolved achiral silicate molecules).

### 6.2. Achiral Minerals with Chiral Faces

Depending on the environmental conditions, minerals can form different crystal habits, some of which may have faces that are chiral [185] (Table 4). For example, calcite ( $\text{CaCO}_3$ ) is centrosymmetric (i.e., its unit cell is achiral), but it can take on many different forms, the most abundant being scalenohedral (e.g., dog-tooth spar) or rhombohedral; the former has chiral faces  $\{2\ 1\ 4\}$  while the latter has faces  $\{1\ 0\ 4\}$  that are achiral [186] (Figure 3). The scalenohedral form comes in pairs of chiral faces, but the overall unit cell remains achiral. This is notably distinct from chiral minerals whose faces are not only all chiral, but are either entirely all L or D forms (rather than having a crystal with both), except for crystals which are internally twinned via the Brazil law. Crystal twinning occurs when a crystal is

subjected to an environmental change (usually temperature or pressure) that results in the growth of a different crystal onto a face of the first. This is especially common for minerals that have polymorphs that are favored under different geochemical conditions. There are different types of twinning; for chiral minerals, the most relevant is the Brazil law twin, which combines right- and left-handed crystals within a growing crystal structure. For quartz, Brazil law twinning is common and is almost never reflected in the surface crystal morphology; thus, quartz minerals may appear chiral, but are internally twinned and are thus racemic [187]. Given the potential for an achiral mineral to possess a racemic interior, together with the fact that achiral crystals can possess a set of chiral faces, it has been advocated that chirality studies involving minerals should avoid powdering the mineral of interest (as chirality information is then lost) [188]. Instead, studies should focus on organic adsorption of chiral species on preserved faces of crystals, whose chirality can be checked against the resulting ee (if any).

### 6.3. Local Chiral Sites

In the natural world, minerals usually form imperfect crystals; that is, their internal structure does not perfectly match that expected from the unit cell. The interruption of the ideal crystal lattice (i.e., crystal defects) can yield local chiral sites on an otherwise achiral mineral or crystal face [185]. Crystal defects are categorized into three groups: point (e.g., vacancies, irregular placement of atoms, substitutions, or kinks), linear (e.g., steps), and plane (e.g., terraces). As organics adsorb most efficiently at defect sites (ledges, kinks, and surface vacancies > terraces), these local chiral centers could provide an efficient means to generate local regions of organic ee [186]. Structural defects are an inherent part of crystal growth; thus, if organics bind to the growing mineral during this process, they can have an impact on the final crystal structure, including the chirality of the mineral [28,182,189]. As every crystal structure will have some crystal defects, local mineral chiral sites may play a bigger role in influencing organic chirality than the inherent chirality of minerals as a whole. In addition to promoting organic adsorption, defect sites are also reactive [190] and can thus serve as an efficient site for further organic reactivity involving the preferentially adsorbed enantiomer.

In addition to local chiral sites existing at mineral surfaces due to structural defects, chiral selectivity can be induced as a result of cooperative effects between chiral organics and the (in)organic analytes that have previously adsorbed at the mineral surface or intercalated (i.e., inserted) between mineral layers. For example, the initial adsorption of one organic enantiomer at the surface could bind in such a way that its orientation further promoted the adsorption of that particular enantiomer. On the other hand, intercalation of other ions/organics could lead to chiral orientations between mineral layers. There are several mechanisms that can induce chirality between layers, including partial ion-exchange (e.g.,  $\text{Al}^{3+}$  for  $\text{Mg}^{2+}$ ) which could lead to multiple sites between any mineral layers with different ions that are of various size and possess different binding capabilities for organic enantiomers; alternatively, ion-exchange could also result in site vacancies that are chiral [191]. Another means of inducing chiral selectivity between mineral sheets is due to differences in the binding orientation and exchange kinetics between the mineral layers and organic enantiomers [192,193]. For instance, a study on the intercalation of L- and D- histidine within vermiculite clay found that the two enantiomers had different effects on the osmotic swelling and d-spacing between the mineral layers [194,195].

### 6.4. Factors That Influence a Mineral's Degree of Enantioselectivity

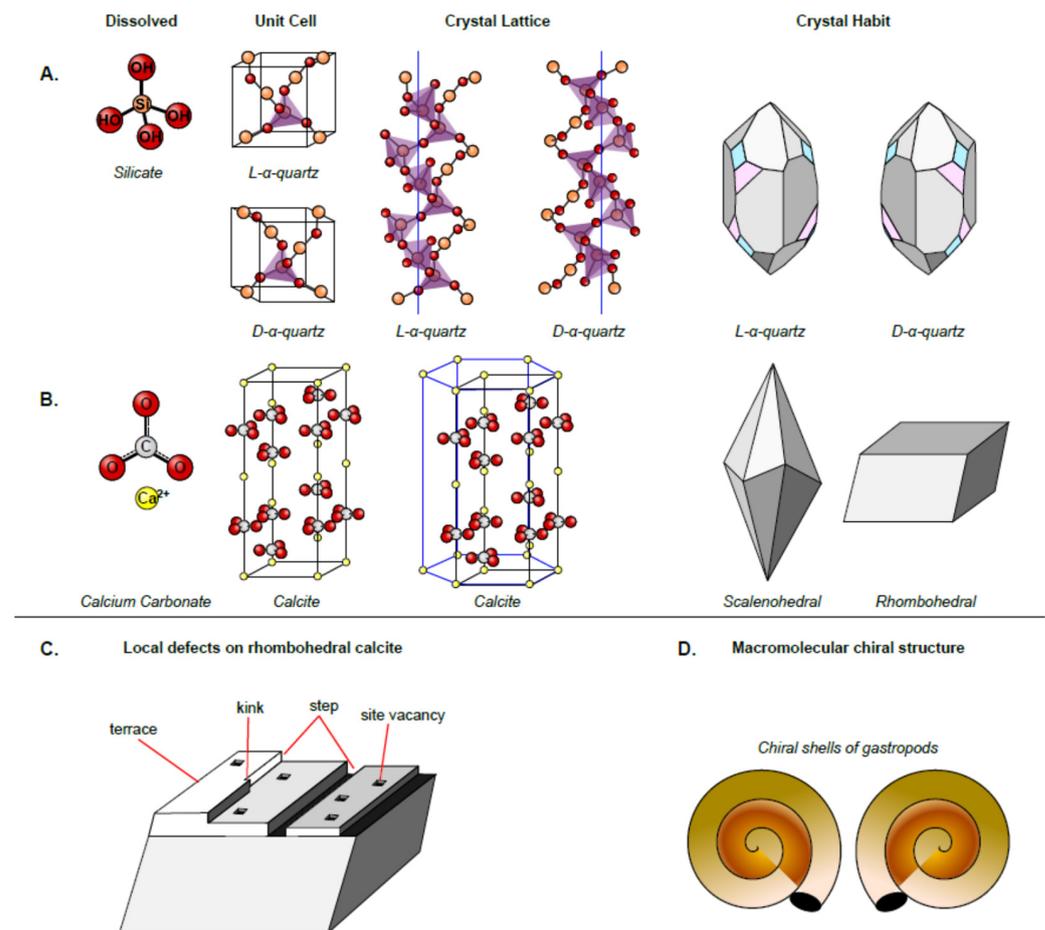
There are inherent factors that influence the degree to which chiral minerals (either via their internal structure or having chiral faces) can drive organic ee: the most notable being the degree to which the mineral enantiomers are structurally different (i.e., the chiral index; [196]) and the degree to which organic enantiomers are sorbed to the mineral surface e.g., ref. [197].

**Table 3.** A list of chiral minerals organized by their crystal system, class, and space group.

Crystal Family	Crystal Class Number	Crystal Class	Space Group	Example Mineral	Formula	Category
Triclinic	1	Pedial	P1	Kaolinite	$\text{Al}_2(\text{Si}_2\text{O}_5)(\text{OH})_4$	Phyllosilicate
				Amesite	$\text{Mg}_2\text{Al}_2\text{SiO}_5(\text{OH})_4$	Phyllosilicate
				Nordstrandite	$\text{Al}(\text{OH})_3$	Metal oxide
Monoclinic	2	Sphenoidal	P2, P2 <sub>1</sub> , C2	Buddingtonite Bassanite	$\text{NH}_4\text{AlSi}_3\text{O}_8$ $\text{Ca}(\text{SO}_4) \cdot 0.5\text{H}_2\text{O}$	Tectosilicate Sulfate
Orthorhombic	222	Rhombic-disphenoidal	P222, P222 <sub>1</sub> , P2 <sub>1</sub> 2 <sub>1</sub> 2, P2 <sub>1</sub> 2 <sub>1</sub> 2 <sub>1</sub> , C222, C222 <sub>1</sub> , F222, I222, I2 <sub>1</sub> 2 <sub>1</sub> 2 <sub>1</sub>	Wülfingite	$\text{Zn}(\text{OH})_2$	Metal oxide
				Epsomite	$\text{MgSO}_4 \cdot 7\text{H}_2\text{O}$	Sulfate
				Sanderite	$\text{MgSO}_4 \cdot 2\text{H}_2\text{O}$	Sulfate
				Lecontite	$(\text{NH}_4, \text{K})\text{NaSO}_4 \cdot 2\text{H}_2\text{O}$	Sulfate
				Abuqite	$\text{CaAl}_2(\text{PO}_4)_2\text{F}_2$	Phosphate
Tetragonal	4	Tetragonal-pyramidal	P4, P4 <sub>1</sub> , P4 <sub>2</sub> , P4 <sub>3</sub> , I4, I4 <sub>1</sub>			
	422	Tetragonal-trapezoidal	P422, P42 <sub>1</sub> 2, P4 <sub>1</sub> 22, P4 <sub>1</sub> 2 <sub>1</sub> 2, P4 <sub>2</sub> 22, P4 <sub>2</sub> 2 <sub>1</sub> 2, P4 <sub>3</sub> 22, P4 <sub>3</sub> 2 <sub>1</sub> 2, I4 <sub>1</sub> 22, I4 <sub>2</sub> 2	Cristobalite Wardite	$\text{SiO}_2$ $\text{NaAl}_3(\text{PO}_4)_2(\text{OH})_4 \cdot 2(\text{H}_2\text{O})$	Tectosilicate Phosphate
Hexagonal	3	Trigonal-pyramidal	P3, P3 <sub>1</sub> , P3 <sub>2</sub> , R3	Monohydrocalcite	$\text{CaCO}_3 \cdot \text{H}_2\text{O}$	Carbonate
	32	Trigonal-trapezohedral	P312, P3 <sub>1</sub> 12, P3 <sub>2</sub> 12c, P3 <sub>2</sub> 12, P3 <sub>1</sub> 21, P3 <sub>2</sub> 21, R32	Berlinite	$\text{AlPO}_4$	Phosphate
				$\alpha$ -D-quartz	$\text{SiO}_2$	Tectosilicate
				$\alpha$ -L-quartz	$\text{SiO}_2$	Tectosilicate
				Antarcticite	$\text{CaCl}_2$	Chloride
				Huntite	$\cdot 6\text{H}_2\text{O}$	Carbonate
					$\text{Mg}_3\text{Ca}(\text{CO}_3)_4$	
6	Hexagonal-pyramidal	P6, P6 <sub>1</sub> , P6 <sub>2</sub> , P6 <sub>3</sub> , P6 <sub>4</sub> , P6 <sub>5</sub>	Trinepheline Kellyite Nagelschmidite	$\text{SiO}_2$ $\text{SiO}_2$ $\text{KAlSiO}_4$ $\text{Mg}_4\text{Al}_2(\text{OH})_{12}(\text{CO}_3) \cdot 3\text{H}_2\text{O}$	Silicate Phyllosilicate Neosilicate	
622	Hexagonal-trapezohedral	P622, P6 <sub>1</sub> 22, P6 <sub>2</sub> 22, P6 <sub>3</sub> 22, P6 <sub>4</sub> 22, P6 <sub>5</sub> 22	$\beta$ -D-quartz	$(\text{Ni,Fe})_4\text{P}$	Tectosilicate	
			$\beta$ -L-quartz	$\text{MnSi}$	Tectosilicate	
			Kalsilite	$\text{FeSi}$	Kalsilite	
			Quintinite	$\text{K}_2\text{Mg}_2(\text{SO}_4)_3$	Carbonate	
Cubic	23	Tetartoidal	P23, P2 <sub>1</sub> 3, F23, I23, I2 <sub>1</sub> 3	Melliniite Brownleeite Naquite	$\text{NH}_4\text{Cl}_y\text{-Fe}_2\text{O}_3$	Phosphide Silicide Silicide
	432	Gyroidal	P423, P4 <sub>2</sub> 32, P4 <sub>3</sub> 32, P4 <sub>1</sub> 32, F432, F432, I432, I4 <sub>1</sub> 32	Langbeinite		Sulfate
				Salammoniac		Chloride
				Maghemite		Metal oxide

**Table 4.** A list of some achiral minerals with chiral faces.

Mineral	Formula	Face {Miller Index}	Category
Calcite	$\text{CaCO}_3$	(214)	Carbonate
Gypsum	$\text{CaSO}_4 \cdot 2\text{H}_2\text{O}$	(110), (111)	Sulfate
Olivine	$(\text{Mg}^{2+}, \text{Fe}^{2+})_2\text{SiO}_4$	(111)	Silicate
Clinopyroxene	$(\text{Ca}, \text{Mg}, \text{Fe}, \text{Na})(\text{Mg}, \text{Fe}, \text{Al})(\text{Si}, \text{Al})_2\text{O}_6$	(110), (111)	Oxide
Clinoamphibole: e.g., hornblende	$(\text{Ca}, \text{Na})_{2-3}(\text{Mg}, \text{Fe}, \text{Al})_5(\text{Al}, \text{Si})_8\text{O}_{22}(\text{OH}, \text{F})_2$	(110), (011)	Inosilicate



**Figure 3.** Summary of the variations of chirality of mineral crystals. **(A)** Inherent chirality, as observed in the unit cell with quartz as an example. **(B)** Chiral faces of achiral minerals with calcite as the example: the unit cell is achiral as is the crystal lattice; however, the crystal habit can have chiral faces, as observed with the scalenohedral habit, and the rhombohedral habit has no chiral faces. **(C)** Natural minerals have local defects that can result in local chiral centers. Shown are defects typical of a growing crystal structure. **(D)** Macromolecular mineralogical structures can have chirality. Often, precipitated shells of biological organisms have a chiral structure, such as the gastropod shells drawn here.

A chiral index: The degree to which enantiomorphic minerals differ from their mirror counterpart is predicted to impact their organic enantioselective potential and can be described mathematically using the chiral index (a term coined by [196]). Downs and Hazen 2004 proposed a methodology to calculate the chiral index of any mineral, and suggested that this measurement could be used to identify minerals with strong potential for driving organic ee. The basis of their calculations follows the logic used to quantify a similar measurement, the distortion index [198], which quantifies the deviation of a periodic arrangement of atoms from its perfect crystal structure due to imperfections. The chiral

index, instead, quantifies the degree of misfit between any two chiral faces or structures. This measurement predicts the potential for mineral faces (e.g., those of calcite) or inherent structure (e.g., quartz) to drive organic ee via mineral–organic interactions. Using this approach, ref. [196] compared the enantioselective potential of the faces of several minerals (calcite, diopside, orthoclase, quartz, and copper) and discovered that the calcite {2 1 4} face, and to a lesser extent diopside and copper, had a higher chiral index (and thus larger enantioselective potential) than the mineral with inherent chiral symmetry—quartz. Although this was a computational study, their results are consistent with experimental work that has found that the calcite {2 1 4} face results in larger ee from racemic amino acid mixtures (~10% chiral excess) than quartz (~1% chiral excess; [187,196]). Note that earlier studies may have underestimated the degree of chiral selection on quartz surfaces due to internal Brazil twinning that results in internally racemic quartz crystals [187,199]; although, studies that took care to work with untwinned quartz crystals still found that the degree of chiral selectivity on quartz is modest at best [200–202].

Downs and Hazen (2004) note that their described chiral index does not necessarily capture the degree to which each mineral can drive organic ee (as there are many factors that can impact mineral–organic interactions). Thus, they suggested two additional chiral indices to consider: one that describes the degree of fit between organic enantiomers onto a mineral surface, or a more general approach based on three-point interactions calculated using triangles formed by nearest-neighbor atoms on the mineral surface. To our knowledge, a follow-up study to calculate these proposed indices has not been published, nor have there been any efforts to apply the chiral index as previously described [196] to other chiral minerals.

**Mineral–organic interactions:** Chiral organic molecules have mostly identical physical and chemical properties; however, the difference in their structural arrangement imparts differences in their degree of interaction with any chiral selector, whether it be organic (e.g., proteins) or inorganic (e.g., minerals). Chiral selectors are inorganic or organic materials that form diastereoisomeric complexes with the chiral analyte. Notably, not all chiral organic molecules will form diastereoisomeric complexes with chiral surfaces on minerals; thus, chiral materials are not always enantioselective. Previous work has demonstrated that there must be a minimum of three noncollinear points of interaction between a chiral analyte and chiral selector for there to be any enantioselectivity; this is also referred to as the three-point minimum interaction model [203–205]. This model is consistent with experimental studies, which found that chiral surfaces of calcite are enantioselective for only certain amino acids (e.g., aspartate), but not others (e.g., alanine, valine, lysine) [186,188,206]. Accordingly, computational studies modeling the interactions of alanine and aspartate with calcite surfaces revealed that aspartate was bound to calcite at three binding sites whereas alanine was only bound at two. While the three-point model is a good starting point for assessing the potential enantioselectivity in the binding of molecules, it should be noted that the model has faced scrutiny [207,208], as it has been demonstrated that sometimes four points of interaction are needed [208]; alternatively, with aromatic organic molecules, a pseudo-two-point interaction may suffice [207].

If we were to develop an equation that properly predicts the enantioselective potential of a mineral in any given environment, the chiral index [196], which describes the mismatch between any two surface enantiomorphs of a mineral, would only represent one variable. Some of the remaining variables would have to capture some information regarding the chiral analyte of interest, including how it interacts with the mineral surface. As previously discussed, generally three points of interaction are required between the mineral and organic analyte for the mineral to impart chiral selectivity onto the system. At these three points, any molecular interaction will suffice, although the type of molecular interaction will impact the degree to which the mineral can drive organic ee and under which environmental conditions it can do so. Mineral–organic interactions can involve bonding and/or nonbonding interactions (i.e., covalent, H-bonds, steric hindrance, pi-pi, ion-dipole, dipole-dipole, dipole-induced-dipole, and London dispersion or van der Waals,

see [209] for a review). The relative strength of these interactions (which is dependent on type and environment) impacts the enantioselectivity of the mineral of interest (with stronger bonds increasing enantioselectivity). Thus, an idealized equation for describing the enantioselectivity of a mineral would also include variables that describe the bond type between each organic–mineral interaction, the surrounding environmental conditions (pH, temperature, and salinity), and how the organic analyte and mineral surface are expected to change with environmental parameters (e.g.,  $pK_a$  of the organic, point of zero charge for the mineral).

#### 6.5. Mineral–Organic Interactions for Driving Enantiomeric Excess

Minerals can drive enantiomeric excess (ee) via selective adsorption that facilitates (1) the production of one enantiomer of a chiral compound; (2) the formation of homochiral polymers (e.g., by facilitating polymerization of L-amino acids); (3) solutions with ee (e.g., preferential adsorption of L-sugars generates solutions enriched in R-sugars); and/or (4) the preferential preservation (or degradation) of the adsorbed enantiomer. Notably, the number of studies that have investigated the influence of chiral mineral surfaces on driving ee of organic enantiomers remains limited. In general, these studies have found that minerals only induce minimal to modest enantiomeric selectivity [177,186,200,201,210]. Although this may result in significant ee over geologic timescales, as chiral molecules can racemize over time, this may not provide an adequate solution ([211] and references therein). One potential resolution to this conundrum are autocatalytic processes, wherein mineral–organic interactions yield conditions that promote enantioselectivity, and the resulting products in turn further promote enantioselectivity. One of the most notable autocatalytic reactions in organic synthesis is the Soai reaction (Table 1), which produces a near enantiopure solution (>99.5% ee) in high yields (>99%) using 3-pyridylalkanol as an asymmetric autocatalyst [26]. As summarized in the review by [212], the organic autocatalyst of the Soai reaction can be successfully substituted for more prebiotically relevant catalysts, including the minerals quartz, gypsum, retgersite, cinnabar, sodium chlorate, and sodium borate—all of which generated near enantiopure solutions in high yields [27,213–217]. These results demonstrate how minerals can induce significant enantioselectivity when subjected to asymmetric autocatalytic reactions.

Autocatalytic systems can also include the precipitation of enantiopure minerals and/or organic crystals as well as formation of surface defects that further promote mineral–organic enantiospecific reactions. These reactions warrant further studies, given that minor ee of either the organic or mineral reagent could lead to near enantiopure solutions or crystals. In fact, even racemic or achiral solutions can produce enantiopure crystals. This was observed in [218], which found that stirred solutions of sodium chlorate (which is achiral) precipitated chiral crystals, all of which had the same handedness; repeated experiments found that the system would consistently produce crystals of either D or L handedness, although which enantiomorph (D or L) precipitated varied between reactions. Subsequent studies elucidated the mechanism at work: the stirring of the solution promoted rapid shearing of the initial crystal that precipitated [15,219]. Thus, if the initial crystal precipitate was of the D form, then shearing (due to mixing) would rapidly break apart the D crystal, producing many smaller D crystals that could then act as nucleation centers for further precipitation of the D enantiomorph. For such reactions, an initial enantiomeric seed was not added to the system, but was a product of random chance.

Although systems that generate enantiopure products from achiral mixtures may have been relevant for the origins of homochiral polymers in biochemistry, the fact that they generate either D or L forms only by chance suggests that overall such systems may have canceled each other out. Alternatively, systems containing a slight excess of an organic and/or mineral enantiomer may, via positive feedback reactions, result in systems with significant ee. One example of such chemistry is that of D and L precipitating tyrosine [220]. Originally, the authors found that D-tyrosine crystallized out of solution much faster than L-tyrosine. These results were replicated by [221] when they used D/L-tyrosine

from the same source as [220]; however, when different tyrosine sources were used, the results were not replicated. Accordingly, when [222] repeated the study, they found that D/L-tyrosine had different solubilities only when the prepared solutions were not passed through a 0.2  $\mu\text{m}$  filter; when the solutions were filtered, the effect was not observed. Both [221] and [222] concluded that the discrepancy of D/L-tyrosine solubility was due to different contaminants, such as fungal spores being present in the D/L-tyrosine powders purchased from various vendors. The authors of [223] contested this interpretation, arguing that the solubility differences were due to the parity violation energy difference between enantiomers (a topic which falls outside of the scope of this review).

If trace contaminants are responsible for the apparent differences in enantiomeric solubility of tyrosine, then a slight excess of a mineral and/or organic enantiomer could drive the precipitation of chiral structures, such as D-tyrosine organic crystals. Alternatively, minor seeds of an organic enantiomer could result in the precipitation of chiral mineral structures, which in turn facilitate the adsorption of an organic enantiomer. For example, D/L enantiomers of aspartic acid or glutamic acid present in solutions undergoing  $\text{CaCO}_3$  precipitation, induced the formation of macro mineral–organic chiral structures [182]. The resulting chiral structure could then promote the adsorption of the seed enantiomer. On the other hand, the presence of organic acids with minor ee can promote the dissolution of the chiral mineral surfaces to which it binds, producing step-like features on the surface [224]. These steps can then enhance organic adsorption onto the corresponding chiral surface, consequently promoting enantioselective adsorption onto that surface and driving ee of the remaining solution [186].

#### 6.6. Future Directions

There are many minerals that are inherently chiral (Tables 3 and S1); however, their prevalence on early Earth or other planetary bodies remains understudied and uncertain. Moreover, widespread conditions favoring the formation of a particular mineral enantiomorph seem unlikely. For example, on Earth, the two quartz enantiomorphs are equally abundant and are often found within the same crystal (i.e., the crystal is internally twinned and thus racemic) [225,226]. Investigations into conditions that could potentially favor the crystallization or preservation of one mineral crystal enantiomorph, over the other, remains uncertain and largely unexplored. One potential research avenue to explore this question would be to computationally explore whether dissolved achiral contaminants ever preferentially adsorb onto only one of two chiral mineral surfaces, thus promoting the dissolution or precipitation of a particular chiral surface. Alternatively, circularly polarized light or spin polarized electrons emitted during beta decay of radioactive nuclei, both of which are proposed as a mechanism for ee of L-amino acids in meteorites [227–229], see 14 and references therein, could also impart chirality onto precipitating minerals or preferentially alter the surface of achiral or chiral minerals, producing surface defects that could either serve as additional chiral sites and/or further promote organic adsorption and reactivity [186,190,230–235]. These astrophysical processes could thus serve as a means for generating inorganic chiral seeds that could subsequently trigger asymmetric autocatalytic systems, as described in Section 6.5.

Prebiotic investigations of the stereoselective effects minerals impart on organic systems and vice versa is relatively limited, with most organic–mineral studies focusing on the adsorption of amino acids on montmorillonite or other clays see [236] and references therein. However, as shown in Tables 3, 4 and S1, there are many minerals that are either inherently chiral or possess chiral faces that are geologically or chemically relevant; even achiral minerals with achiral faces can possess local chiral sites at the surface or between layers. The most notable minerals with chiral surfaces are those of evaporites, which are especially interesting considering their prevalence on the Martian surface [237,238]. Minerals with inherent chirality and planetary relevance include some phyllosilicates (e.g., kaolinite, a common crustal mineral identified on Mars [239–241]), zeolites (which can act as prebiotic catalysts and has been detected on Mars, [242–245]), sulfates (e.g., epsomite,

bassanite, and sanderite, which are present on the Martian surface and can preserve biosignatures [238,240,246–248]), sal-ammoniac ( $\text{NH}_4\text{Cl}$ ), which has been found on the surface of Ceres [249,250], maghemite (a common Fe oxyhydroxide that can form from weathering of magnetite and is also likely on Mars [251–254]), phosphates (which could act as a source for organic phosphorylation of chiral compounds), and borates (which are the likely form of boron detected on the Martian surface and have been demonstrated to selectively promote the preservation of ribose [255–260]). Given the planetary and prebiotic relevance of such chiral minerals, additional investigations exploring their enantioselective potential within prebiotic systems are warranted.

As previously discussed, the enantioselective potential of any chiral surface is dependent on its local environment (pH, temperature, salinity, and pressure), the organic analyte ( $\text{pK}_a$ , and number and types of bonds with the mineral), and its inherent physical properties (point of zero valent charge and degree of mismatch between its two enantiomorphs). The authors of [196] calculated a chiral index for several chiral minerals based on the degree of mismatch between their enantiomorphs, and suggested that other chiral indices could be calculated based on the three-point interaction model and the differences between the binding sites of two organic enantiomers onto any chiral surface. This original study was published in 2004, and since then molecular modeling approaches have vastly improved and now provide a facile means to obtain the geometry of the available binding sites on any mineral surface and the resulting geometry and bond energies for each of the mineral–organic interactions. These variables could then be used as inputs to calculate a better description of the enantioselective potential of a wide-range of minerals under various geochemical conditions. Such studies are warranted, as obtaining enantiopure minerals is challenging (as they may look enantiopure on the outside, but possess internal twinning) and that natural minerals also possess organic contaminants, many of which may be biological and thus have their own asymmetry that could influence the reaction system being studied. In sum, molecular modeling can be a boon to prebiotic studies of ee, as it can be leveraged to systematically evaluate a wide range of minerals and organic–mineral interactions under various geochemical conditions and thus help shed light onto which chiral minerals and mineral surfaces warrant further experimental study.

## 7. Alteration during Geochemical Processes

### 7.1. Terrestrial Geochemical Alteration and the Preservation of Organics

If one were to order the classes of organics based on preservation potential and biological specificity, there would be an inverse relationship [261]. While DNA, RNA, and proteins can specifically fingerprint their biological source, they are also the most labile organic compound class [262,263]. Lipids, on the other hand, while present in many living organisms and can have high biological specificity, typically exhibit a broad range of sources. Lipids and their hydrocarbon derivatives are the most recalcitrant class of organic compound; therefore, over geologic time and under suitable conditions, they will be the most likely organic remnants [264,265].

There are some asymmetric reactions that are involved in the synthesis of (phospho) lipid membrane molecules (e.g., [266–269]). Although lipids are considered the most taphonomically robust to degradation and alteration over long geologic timescales, they are not immediately associated with research into chiral organic compounds. Lipids possess chirality in the way that amino acids and sugars do, and commonly have multiple stereogenic centers leading to many different potential stereoisomers (discussed below). To understand how organics and biosignatures can be detected in samples, it is important to recognize the reactions and conditions that affect their sequestration and preservation. The following sections will describe the processes by which organic compounds are sequestered and preserved in the rock record and the ways in which chiral configurations can be affected by geological alteration.

### 7.2. Formation of Insoluble Macromolecular Organic Matter

On Earth, an average of >99% of organic matter (OM) produced from photosynthesis is rapidly remineralized during early diagenesis by biochemical degradation [270]. As OM is sequestered and buried in the sedimentary record, the physical and chemical processes known as diagenesis, catagenesis, and metagenesis take place. Diagenesis occurs during shallow burial at low temperatures, and physical and chemical processes continue to accumulate a small fraction of degradation-resistant OM, which is polymerized and crosslinked to form a biomacromolecular matrix, known as kerogen. During this process, many lipid components, which can act as biosignatures and are highly resistant to biochemical degradation, are bound within the insoluble organic matter (IOM) matrix [271]. Despite the rapid overturn of OM during the early stages of deposition [272] and the specific environmental conditions required for organic preservation [273], kerogen-like macromolecular OM is recovered even during this early phase [274,275] with the proportion of organics bound into the matrix increasing over geological time [275–277]. Catagenesis involves increases in burial temperature and depth over time with a significant component of heating that induces the cracking of the macromolecular structure to liberate shorter-chain hydrocarbons [278]. Metagenesis occurs at higher burial temperatures (low grade metamorphic temperatures above 200 °C) and depths and involves the cracking of residual hydrocarbon bitumen into dry gas (>98% methane) and solidified pyrobitumen [273,279].

Although the definition of kerogen has evolved over time, traditionally it is operationally defined as the solid residue that remains after OM has been extracted with organic solvents [280–282]. For mature, consolidated kerogens, it typically also excludes the acid-digestible and hydrolyzable component containing organics that are associated with the mineral matrix [281]. The remainder of sedimentary OM is defined as the component that is readily extractable with organic solvents, known as bitumen. The components of kerogen consist of the sum of the simple organic precursor molecules that were sequestered into the macromolecular structure during sedimentary diagenetic and catagenetic processes.

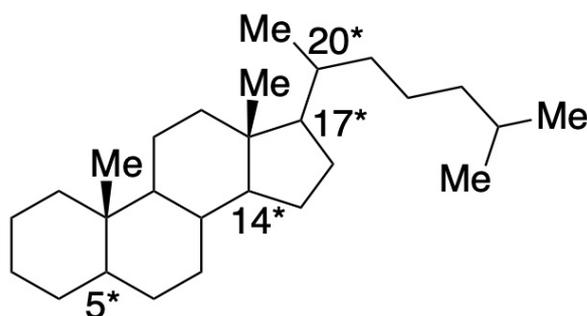
Mineral substrates and matrices are critically important to the preservation of organic compounds [282–284], as they provide stability and protection to organics not only during deposition and sedimentary alteration, but also against harsh environmental conditions, such as radiation (e.g., [285]). The kerogen macromolecular matrix provides long-term preservation of organics by protecting bound molecules from oxidation and degradation [286]. Kerogen is considered an immobile solid matrix; molecules linked within it were deposited synchronously with the host sedimentary rock. Bound lipids released from kerogen are much less prone to organic contamination from migrating fluids relative to their solvent extractable counterparts [287].

Kerogen is the largest sink of organic carbon on Earth [288] and expected to be the dominant mass fraction of preserved organic matter on other planetary bodies [21,289]. Post-depositional alteration processes, such as the binding of organics, are affected by steric hindrance, especially as the kerogen evolves and becomes more structurally and chemically complex. Hence, stereochemistry is an important consideration for the studies of bound and released organics from mineral and organic matrices. Utilizing techniques that release organic constituents either thermolytically (e.g., heating through pyrolysis) or chemolytically (e.g., selective chemical degradation through chemolysis), significantly enhances the analysis of kerogen, as both processes release products that are amenable to standard chromatographic resolution [287].

### 7.3. Stereochemistry of Lipids in Kerogen

The stereochemistry of the organic compounds, specifically the ratio of isomers that are released through various experimental methods, can reveal the thermal maturity conditions during sequestration. An example comes from common microbial cellular membrane lipids, steroids and hopanoids, which are polycyclic biomarker precursors derived primarily from eukaryotes and bacteria, respectively [278]. In the specific case of steroids, the C<sub>27</sub> steroid

(cholesterol) possesses eight asymmetric carbons in its structure, at C-3, 8, 9, 10, 13, 14, 17, and 20 [290], giving rise to the possibility of numerous stereochemical configurations (Figure 4).



**Figure 4.** Structure of the  $C_{27}$  sterane, cholestane. The asterisk (\*) at C-5, C-14, C-17, and C-20 indicate where a H can be in the  $\alpha$ - or  $\beta$ -configuration, and the carbon at C-20 can be the R or S enantiomer. The R enantiomer is the biological form, whereas the S enantiomer is the geologically stable form.

However, only one biological isomer of cholesterol exists, due to the high specificity of steroid biosynthesis [291]. The stereochemical configuration of immature/biologically-inherited steroids and hopanoids as their hydrocarbon equivalents are 20R-5 $\alpha$ / $\beta$ (H), 14 $\alpha$ (H), 17 $\alpha$ (H)-cholestane ( $C_{27}$  sterane), and 17 $\beta$ (H), 21 $\beta$ (H)-hopane ( $C_{30}$  hopane), respectively. These stereochemical configurations are thermally unstable and they isomerize to the stable forms of 20S-5 $\alpha$ (H), 14 $\beta$ (H), 17 $\beta$ (H)-cholestane for steranes and 17 $\alpha$ (H), 21 $\beta$ (H)-hopane for hopanes over geological time, as these molecules relieve ring strain and convert from the flat, planar configuration to a bent configuration [278]. During incorporation into kerogen, the asymmetric centers of steranes and hopanes are protected from further thermal isomerization due to steric hindrance within the matrix [278], although the stereospecificity that is highly conserved in the biological form is often lost, with an array of isomers forming [292].

Isomerization, while primarily due to increases in temperature through thermal alteration, could also be catalyzed by clay minerals [293,294]. It has been demonstrated that mineralogy and salinity can affect oil generation kinetics as well as product composition [295,296]. Experimental techniques for fragmenting kerogen into analyzable components, such as pyrolysis assisted by high-pressure hydrogen gas (hydrolysis; [287]), releases bound lipids that conserve the original stereochemistry during sequestration [275,297–301]. Thermochemolysis using tetramethylammonium hydroxide (TMAH) compared to standard hydrous pyrolysis demonstrates differences in the stereochemical configuration of released hopanoid products [302], indicating that configurational isomerization (epimerization) occurs during the analytical processes. Future experiments that track structural modifications to stereochemistry associated with simulated alteration and/or maturation processes will be critical to understanding the underlying mechanisms and transformation pathways that prevail during natural processes. They will help better identify and optimize the experimental techniques that should be employed to study these chemical reactions.

#### 7.4. Compositional Alteration on Other Planetary Bodies

Meteorites can be classified as chondrites that form from protoplanetary disk material [303,304] and achondrites, which have undergone secondary processes, such as melting and differentiation [305,306]. Chondrite meteorites can be separated into ordinary, carbonaceous, Rumuruti, Kakangari, and enstatite chondrites [304,307]. The carbonaceous chondrite group can be further subdivided into eight main groups (Table 5), based on chemical (mineralogy and isotopic) composition and degree of alteration [308,309]. The six petrologic types describe the extent of aqueous (type 1, 2, and 3) and thermal (types 3, 4, 5, and 6) alteration, where type 1 is the most aqueously altered and type 6 is the most thermally altered [310]. Similar to how alteration and structural modifications to chemical

composition can occur to organic matter on Earth, alteration can and does occur during the preservation of carbonaceous OM on other planetary bodies, such as asteroids, comets, and planetary surfaces. These matrices are typically minerals associated with rocky or icy substrates and can occur on the surface regolith or in the parent body interior [311]. Alteration can occur through thermal or aqueous processes, shock wave propagation, and brecciation.

**Table 5.** Properties of carbonaceous chondrite meteorite groups, including: matrix abundances, chondrule abundances and sizes, refractory component abundances, metallic Fe and Ni abundances, average olivine compositions, and refractory lithophile element abundances. The carbonaceous chondrite meteorite groups are arranged from left to right (CI to CB) in order of decreasing bulk rock oxidation. Table is from [14] and data were compiled from [307,312–314].

Petrologic Type	CI	CM	CK	CV	CO	CR	CH	CB
Petrologic type	1	1–2	3–6	2–3	3	1–2	3	3
Chondrule abundance (vol.%)	≤1 <sup>†</sup>	20 <sup>‡</sup>	15	45	40–48	50–60	~70	20–40
Matrix abundance (vol.%)	>99 <sup>†</sup>	70 <sup>‡</sup>	75	40	30–34	30–50	5	<5
Refractory abundance <sup>#</sup> (vol.%)	≤1	5	4	10	13	0.5	0.1	<0.1
Metal (Fe,Ni) abundance (vol.%)	≤1	0.1	≤1	0–5	1–5	5–8	20	60–80
Average chondrule diameter (mm)	n.a.	0.3	0.7–0.8	1.0	0.15	0.7	0.02–0.09	0.2–10
Olivine composition (mol% Fe <sub>2</sub> SiO <sub>4</sub> ; range)	*	*	<1–47	*	*		<1–36	2–3
(mol% Fe <sub>2</sub> SiO <sub>4</sub> ; mode)			29–33			1–3	2	3
Refractory lithophiles <sup>#</sup>	1.00	1.15	1.21	1.35	1.13	1.03	1.00	1.0–1.4

<sup>†</sup> Including chondrule fragments and silicate minerals inferred to be chondrule fragments changes matrix and chondrule abundances to >95 vol.% and <5 vol. %, respectively. <sup>‡</sup> Variable. <sup>#</sup> Calcium aluminum inclusions + amoeboid olivine aggregates. \* Highly variable and unequilibrated. <sup>#</sup> Mean ratio refractory lithophiles relative to Mg and normalized to CI chondrites.

Thermal alteration characterized by increases in temperature can affect the mineralogical composition of chondrites through metamorphic reactions and localized recrystallization but not widespread melting (e.g., [315–319]). CK chondrites are a unique group of carbonaceous chondrites and are the only group that has experienced the full range of thermal alteration types [210,320]. Thermal modeling has been used to constrain the cooling histories of different chondrite parent bodies and is often achieved through the use of elemental gradients, trace elemental and isotopic composition, and radiometric ages calculated from the closure temperature of minerals during crystallization [321–325].

Aqueous alteration occurs when water ice in the parent body melts due to heat generated through radioactive decay of elements [326]. Water can alter the primary composition (both mineralogically and isotopically) of parent body materials through dissolution and formation of secondary hydrous phases, such as clays, serpentines, carbonates, sulfates, sulfides, halides, oxides, and oxyhydroxides [327]. Aqueous alteration can be determined through petrographic analysis of minerals, stable and radiogenic isotopes [328–332], as well as microtextures [333]. Although most chondrite groups have been affected by aqueous alteration to some degree, there are different types of aqueous alteration. Low temperature hydrothermal alteration occurs at temperatures below 200–300 °C [334] and can form mineral phases, such as phyllosilicates. At higher temperatures, typically above 300 °C [335], fluid-assisted metamorphism alters mineral composition. At similar temperatures, but in an open system where fluids are flowing through the rock and altering the composition, is referred to as metasomatism [336]. Some recent experimental studies have suggested that the initial parent body composition need not contain water, as heat from impact [337] or organic degradation [338] could induce aqueous alteration on parent bodies.

Alteration to physical and chemical properties by shock is a common and ongoing phenomenon in chondrites; this includes structural changes due to collision with meteorite parent bodies that can be related to pressure and temperature increases [339]. Shock propagation features can be difficult to deconvolve due to the heterogeneous nature of

chondrites [340]. Changes to composition can include localized melting at the site of impact, as well as fracture features.

Formation of meteorite breccias (brecciation) involves the high-speed impact of meteorites that combine with other fragments of parent bodies [341]. Studying meteorite breccias is important for understanding early Solar System processes and combines a variety of the above-mentioned techniques, such as shock wave propagation, impact velocities, and textural analysis. Due to the variation in composition of the brecciated fragments, multiple parent body sources may be invoked.

Carbonaceous chondrites and the OM contained within them have been subject to thermal and aqueous processing. There are numerous studies that report on the effect of parent body alteration on amino acids, carboxylic acids, and nucleobase compositions [34,40,42,342–346]. The degree to which parent-body processing can affect organic molecule populations is largely dependent on the degree and type of alteration [347]. Petrological type 3 chondrites have undergone little to no thermal or aqueous processing (pristine). Types 1 and 2 have experienced aqueous alteration, while types 4, 5, and 6 have experienced thermal alteration, although not enough to start differentiation. The inventory (specific molecules and abundances) of organics can range markedly across meteorite class and degrees of alteration. For example, [45] reported amino acid distributions in CB and CH chondrites that differed markedly from those reported for type 2 and 3 CM and CR chondrites. The CB and CH chondrites contained beta-, gamma-, and delta-amino acid abundances that were much higher than the relative alpha amino acid abundances, evidencing multiple formation mechanisms. CI chondrites, which are type 1 (highest degree of aqueous alteration), have been reported to contain a higher abundance of beta-alanine relative to alpha-alanine and glycine, which is not the case of CM chondrites [42,348]. This suggests distinct parent bodies for CM and CI chondrites.

## 8. Laboratory Analysis and Reactions

### 8.1. Analytical Instrumentation

In laboratory settings, mineral and metal substrates have been demonstrated to be able to mediate different types of organic reactivity. This includes the synthesis of prebiotic materials, such as nucleobases and amino acids, as well as promoting metabolically relevant reactions [3–6,349–353]. In many of these cases, the enantioselectivity of the reaction was not explored. As we are growing to understand these factors, there are several avenues of research that can be explored to answer questions relevant to chirality at the origins of life. Herein, we summarize the alteration processes of minerals as well as relevant organic reactions that could impart asymmetry. We also make recommendations for future research endeavors in this field.

While the focus of this paper is organic reactions relevant to prebiotic chemistry and the origins of life, Table 1 highlights the diverse applications of enantioselective reactions for pharmaceutical, natural product, and other syntheses. Enantioselective synthesis plays a major role in pharmaceutical research, as drug design and testing investigate biological and pharmacological compounds that are chiral and whose chirality has significant implications on their active properties in biological systems [102,354,355]. As a result, there is much research centered around the analysis and detection of chiral compounds, the foremost being separation of enantiomers using chiral column chromatography [356,357]. Chirality can also be distinguished using spectroscopic and calorimetric techniques, and various other separation strategies (Table 6). While enantiomers have the same nuclear magnetic resonance (NMR) spectra, derivatization could allow for visible differences in the NMR by generating diastereomers or employing the use of chiral solvating agents without the need for diastereomer distinction [358]. Light scattering has been used previously for enantiomeric separation; however, this is extremely sensitive to other materials within the solution [359,360] and is currently not widely used [229].

**Table 6.** Analytical instrumentation used for separation and detection of racemic mixtures into chiral enantiomers grouped by primary analytical techniques. Included are their relevant fields of application and whether or not the technique used for chiral separation has been proposed or demonstrated in spaceflight.

Instrument	Separation	Detector	Application	Mission Relevance	References
Chromatography and Spectrometry					
Gas chromatography–mass spectrometry	GC with a chiral column	MS	Organic chemistry; origins of life	Cometary Sampling and Composition (COSAC)-Rosetta: launched March 2004 but sampling unsuccessful Sample analysis at Mars-Mars Science Laboratory: in progress, landed August 2012 Mars Organic Molecule Analyzer (MOMA)-ExoMars: planned September 2022 launch	[361–365]
Liquid chromatography–mass spectrometry; high performance LC-MS	(HP)LC with a chiral column	MS (various)	Organic chemistry; origins of life	No	[366–368]
Sub- and supercritical fluid chromatography (SFC)	SF (CO <sub>2</sub> plus polar co-solvent)	Various: UV-Vis, diode-array, evaporative light scattering (ELS) detector, charged-aerosol detection, MS (atmospheric pressure chemical ionization, electrospray ionization)	Organic chemistry; forensics	No	[369,370]
Capillary electrophoresis (CE)	CE	Laser-induced fluorescence (LIF)	Origins of life; organic chemistry; instrument development	Proposed	[371–373]
Capillary electrochromatography (CEC)	CE/HPLC	Various; UV detectors	Organic chemistry	No	[374–376]
Ligand exchange CE	CE	Various; UV detectors	Organic chemistry	No	[377–379]
Non-aqueous CE (NACE)	CE	Various detectors; UV, conductivity, MS, LIF	Organic chemistry; medicine	No	[380]
Ion-mobility mass spectrometry (IM-MS)	Derivatization, chiral neutral gases	IM-MS	Organic chemistry; origins of life	Volatile Organic Analyzer (VOA) on the International Space Station (ISS); for air quality control not enantiomeric separation—deployed August 2001	[381–384]

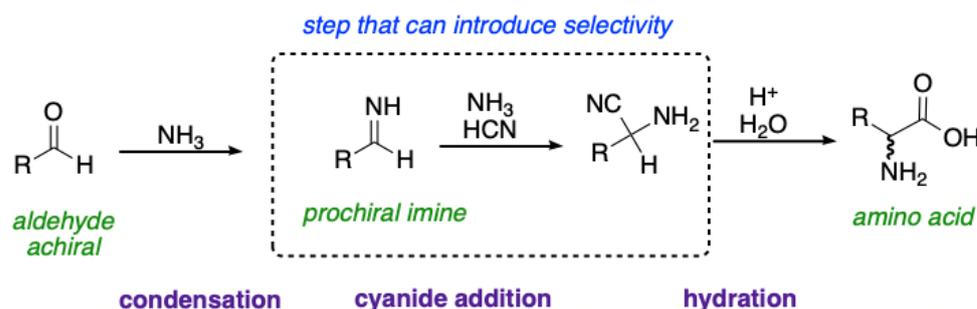
Table 6. Cont.

Instrument	Separation	Detector	Application	Mission Relevance	References
Photodissociation	Photodissociation in cold gas phase	Various MS; e.g., ESI	Biochemistry	No	[385–387]
Matrix-assisted laser desorption ionization (MALDI)—time of flight (TOF) MS	Stereosensitive fragmentation (SF)	MALDI-TOF/TOF MS	Biochemistry	No	[388]
Spectroscopy					
Nuclear Magnetic Resonance (NMR)	Various, derivatization (typically to form diastereomers)	NMR	Organic chemistry	No	[358,389]
Ultraviolet (UV)-visible (Vis) spectrophotometry	Various	UV-Vis	Organic chemistry	No	[390,391]
Infrared (IR) spectroscopy	Various, e.g., CE, NACE	FT-IR	Organic chemistry	No	[392,393]
Optical rotatory dispersion (ORD)	Polarized light	Detector	Organic chemistry	No	[394,395]
Circular dichroism (CD)	Circularly polarized light	CD detector (various)	Organic chemistry; biochemistry	No	[396–398]
Femto-second (fs) laser mass spectrometry	fs-laser	MS	Organic chemistry	No	[399]
Polarimetry	Various; cavity ringdown, near IR	Detector, photodetector	Materials science; origins of life	Proposed	[400–402]
Optical techniques					
Evaporative light scattering (ELS)	Hydrophilic interaction chromatography (HILIC)	Light scattering detector (LSD)	Organic chemistry	No	[403]
ELS	High performance liquid chromatography (HPLC)	LSD	Organic chemistry	No	[404]
Laser	Off-resonant laser beam	Detector	Nanotechnology	No	[405]
Atomic force microscopy (AFM)	Optical tweezers	Optical and AFM	Nanotechnology; materials science	No	[406]

Table 6. Cont.

Instrument	Separation	Detector	Application	Mission Relevance	References
Polarization camera	Micropolarizer array	Detector	Origins of life	Proposed	[407]
Calorimetry					
Differential scanning calorimetry (DSC)	Thermal	Calorimeter	Organic chemistry; macromolecules	No	[254,408]
Separation					
Batch crystallization	Various; e.g., chromatography	Model that calculates the optimal conditions for separation	Organic chemistry	No	[409]
Diastereoisomeric recrystallization	Crystallization	Various, e.g., MS, DSC, X-ray diffraction (XRD)	Organic chemistry	No	[410–412]
Kinetic resolution	Various, e.g., chiral catalysts	Various, e.g., HPLC-MS, ESI-MS	Organic chemistry	No	[413–415]
Labeling					
Fluorescent sensors/dyes	Various dyes, e.g., 5-carboxyfluorescein succinimidyl ester, fluorescamine	Various fluorescence detectors (e.g., confocal fluorescence microscope)	Origins of life; organic chemistry	Proposed	[416–420]

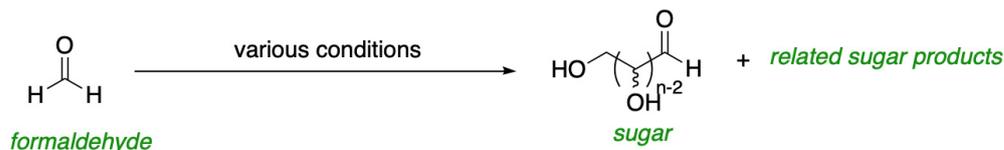




**Figure 6.** Reaction mechanism of an achiral aldehyde in presence of ammonia and cyanide to form an  $\alpha$ -aminonitrile intermediate and an amino acid following hydrolysis.

### 8.2.3. Formose Reaction

The formose reaction begins with formaldehyde, which, once it condenses to form glycolaldehyde, reacts autocatalytically to form sugars abiotically under alkaline conditions (Figure 7; [439]). Unlike the above examples, this reaction forms a variety of products related to sugars, including methanol, formic acid, sugar alcohols, branched sugars, sugar oligomers, sugar acids, and hydroxy acids, to name a few [174,440–442]. The formose reaction can be mediated by many different materials, including divalent cations, such as  $\text{Ca}^{2+}$  [443–446], and silicate minerals [447]. For a detailed review on conditions conducive for the formose reaction, see [422] and references therein.



**Figure 7.** General outline of the formose reaction.

The formose reaction underpins the leading theory for the origin of sugars under prebiotic conditions (although the glyoxylate scenario is a proposed alternative [448]) and as such, further studies of asymmetric formose reactions are therefore particularly important, especially considering that the formose reaction is autocatalytic and could thus generate large ee if mediated by even scarce amounts of a chiral seed (see Section 6.5). Previous work has demonstrated how formose reactions can result in ee of the resulting D-sugars when conducted under conditions, such as UV-irradiation [449], or seeded with low concentrations of L-amino acids, namely proline [450,451].

### 8.3. Solid-State Reactions

While many reactions of interest to the prebiotic chemistry field are in the solution phase (homogenous or heterogenous), solid-state chemical reactions could also be relevant to this field. It has been demonstrated that amino acids, nucleotides, nucleosides, peptides, and sugars can be formed under mechanochemical conditions [452–455]. Metal-mediated and metal-catalyzed reactions have been demonstrated under mechanochemical conditions [456]. There are also examples of asymmetric reactions in the solid state [457,458]. Mechanochemical studies utilize mechanical energy to generate reactivity [459]; common examples include ball mill grinding and milling. The advantages associated with mechanochemical synthesis include increased productivity and efficiency [460], as well as resource conservation. These experiments do not require hazardous or expensive solvents and have overall lower energy requirements, vis-a-vis solution reactions [461–463].

While this is used in organic chemistry and materials science [464,465], it is also applicable to planetary science and prebiotic chemistry as a way to mimic bombardment and other interactions. Early prebiotic organics could have been synthesized endogenously or delivered exogenously through impacts via asteroids or comets [466–469]. There is interest in simulating the effect of impacts on the petrologic [470,471] and organic [472–478]

composition of prebiotically relevant materials, through the mechanochemical action of shock-wave propagation. In addition to simulating cometary impact events, there are impact simulations that investigate the effect of hypervelocity on the breakdown of organic compounds that are relevant for biosignature detection in spaceflight, particularly regarding spacecraft velocity during sampling [479,480].

## 9. Recommendations for Future Research

The origin of homochirality of life is a very broad topic; therefore, we recommend that multiple avenues should be explored. This includes experimental conditions/studies that are not directly applicable to prebiotic chemistry on Earth, which could be useful for understanding abiotic sources of organic homochirality and applied to the interpretation of potential biosignatures. We recommend, in the case of mineral-mediated organic reactivity, that the chirality of both the mineral substrate as well as the organic compound be taken into consideration. There has been a significant amount of research exploring mineral-mediated prebiotic reactions and exploring such reactions with chiral mineral mediators would augment understanding.

### 9.1. Capabilities of Flight-Ready Technology

Significant ee is a potential biosignature [48,481]. As we look to find biosignatures on other planetary bodies, homochirality and organic preservation are two areas that deserve focus. The ability to determine chirality on other planetary bodies is an attractive endeavor, particularly in the astrobiology field. Additionally, organic preservation has been demonstrated to be enhanced by minerals and macromolecular matrices that provide protection against degradation; thus, organic–mineral interactions are also an area of importance when looking for potential biosignatures on other planets. There are a variety of instrument technologies that have been proposed and designed for space missions, which include a strong focus on colocated mineral and organic detection [482,483] and the separation of chiral organic molecules (e.g., [361,363,364]). The following sections describe the instruments, on three currently active missions (as of this writing), which are focused on the detection of chirality as a potential organic biosignatures.

#### 9.1.1. COSAC on Rosetta's Philae Lander

The separation of chiral organics via the chiral column GC-MS has been proposed and flown on missions to comets [361,484,485] and Mars [363,364]. The European Space Agency (ESA)'s Rosetta mission launched a space probe in March 2004 to investigate the comet 67P/Churyumov-Gerasimenko. The mission consisted of the Rosetta spacecraft orbiter and the Philae lander. The ten-instrument suite on the Philae lander was designed to investigate the physical and chemical composition of a cometary nucleus through elemental, isotopic, mineralogical, molecular, surface, subsurface, and structural analysis. The plan was to accomplish these goals via drilling, imagery, spectrometry (gas chromatography and alpha particle X-ray), gas analysis, radio transmission, surface and subsurface sensors, and magnetometry [486].

The chirality experiment involved two instrument subsystems; the sampling system (Sample Drilling and Distribution; SD2) and the GC-TOF-MS (Cometary Sampling and Composition; COSAC). SD2 consisted of a drill that was able to penetrate up to 230 mm into the subsurface to collect samples to deposit into a carousel with 26 ovens. The COSAC GC and TOF-MS had a total of eight GC columns on board, three of which were chiral GC columns (Table 7) with the capability to separate out and analyze amino acid enantiomers. Due to complications that occurred during Philae's detachment from the Rosetta spacecraft and landing on the surface of 67P/Churyumov-Gerasimenko in November 2004, the COSAC experiment was not attempted and samples could not be acquired [365]. Although the chirality experiment was not conducted due to landing issues, COSAC was able to sample the atmosphere in a passive mode, 25 min post-landing. The acquired mass

spectrum indicated 16 organic species were present, including four compounds that had not previously been detected in comets [487].

**Table 7.** Chiral GC columns flown or proposed on missions for enantiomeric separation of organics. ESA—European Space Agency; NASA—National Aeronautics and Space Administration; MSL—Mars Science Laboratory; COSAC—Cometary Sampling and Composition; SAM—Sample Analysis at Mars; MOMA—Mars Organic Molecule Analyzer.

Agency	Mission	Status	Instrument	Total GC Columns	Chiral Column(s)
ESA	Rosetta	Flown but unsuccessful	COSAC	8	Chirasil Dex CB Chirasil L Val Cyclodextrin G-TA
NASA	MSL	In progress	SAM	6	Chirasil- $\beta$ Dex CB
ESA/Roscosmos	ExoMars	Planned	MOMA	4	CP Chirasil Dex

### 9.1.2. SAM on MSL's Curiosity Rover

The Martian surface and subsurface has long been considered a prime astrobiological life detection target [48,488–490]. Currently, information about the organic inventory and geological context of Mars is sourced from the study of Martian meteorites [491–493] as well as landed spacecraft that have or are currently exploring the Martian surface [494,495]. NASA's Mars Science Laboratory (MSL) Curiosity rover landed in Gale crater in August 2012. The instrument payload suite includes the Sample Analysis at Mars (SAM) instrument.

The SAM instrument's primary scientific goals include characterization of the composition of Martian atmosphere and the composition of organics within surface and subsurface sediments [363]. SAM consists of a tunable laser spectrometer (TLS), GC, and a quadrupole mass spectrometer (QMS). The GC has six chemically different 30 m columns for the separation of polar and non-polar organic compounds, one of which is a chiral column intended to separate volatile organic compound enantiomers (Table 7; 363). In a laboratory experiment designed to simulate the flight conditions of the chiral column on the SAM GC, Chirasil- $\beta$  Dex was used to separate out chlorohydrocarbon compounds at low (35 °C) and high (185 °C) temperatures. The temperatures were chosen to approximate the dynamic range of temperatures that would be experienced in flight (30 °C and 200 °C are the minimum and maximum GC operating temperatures on SAM, respectively). The results of the experiment demonstrated that 1,2-dichloropropane enantiomers were partially separated at the low temperature; however, at the higher temperature experiment, the enantiomers coeluted [496].

### 9.1.3. MOMA on ExoMars' Rosalind Franklin Rover

ESA's and Roscosmos's Rosalind Franklin rover, part of the ExoMars mission, is scheduled for launch in September 2022 to land in Oxia Planum, Mars, in June 2023 [497]. The goal of the ExoMars mission is to look for signs of past or present life using an instrument suite that will characterize the geochemical environment contained in the surface and subsurface. With a drill capable of reaching 2 m into the subsurface, samples taken by Rosalind Franklin have a higher likelihood of containing sediments that are less affected by radiation and surface oxidation [498].

The Mars Organic Molecule Analyzer (MOMA) consists of a pyrolysis GC-MS and a laser desorption ionization MS (LDI-MS). These two techniques, which ionize samples thermolytically through pyrolysis and are induced by laser ionization, ensure that a wide range of small organic compounds, as well as larger refractory phases, will volatilize for spectral analysis. The GC portion of MOMA contains four chemically distinct GC columns, one of which is a chiral column (Table 7) capable of separating enantiomers. Portions of the MOMA payload contain hardware heritage that is similar to the COSAC experiment aboard Rosetta and SAM aboard Curiosity [364].

Along with the flight models of the instruments discussed above, there exist many brassboard instruments in laboratories that can be used to enhance our knowledge and extend the capabilities of flight technology but with the full range of resources accessible to ground laboratories on Earth (e.g., [499]). We recommend continued testing of analog instruments that are in flight as well as proposed flight-ready instruments at all technology readiness levels (TRLs), along with the traditional techniques used for chiral detection. The generation of extensive datasets from analog and brassboard instruments will help to calibrate and interpret results returned from missions, as well as optimize parameters to be used in new instrument technologies.

### 9.2. Next Generation Instrumentation

In addition to existing instruments, the development of new instruments or instrument capabilities to detect chirality in spaceflight is a necessary component in the search for signs of past life. Advancements in the fields of organic separation (e.g., [500–504]) and detection through mass spectrometry [505,506] are needed to enhance the field of chiral separation on spaceflight missions. Refinement of traditional chiral separation and detection methods as well as the development of new analytical strategies to investigate chiral organics [407,420,507–509] and minerals [510] should be prioritized. Analytical advancements, coupled with progress in spaceflight technology, represent an exciting step forward in the study of prebiotic chemistry and the detection of chirality on Earth and other planetary bodies.

### 9.3. Contamination Control

Habitability and life detection missions commonly need to address issues related to contamination and outgassing in accordance with guidelines suggested by the International Committee on Space Research (COSPAR) for planetary protection (e.g., [48,511,512]). Missions and proposed mission concepts that address contamination control issues are typically or wholly concerned with adhering to the requirements set out by planetary protection (PP) guidelines, through verification and quantification of bioburden on instrument and spacecraft surfaces. Bioburden, a quantitative measure of the number of viable microorganisms on any given surface, is not only important to verify from the perspective of PP (so as not to contaminate a planetary body with Earth microorganisms and vice versa), but to ensure that the in situ measurements being made in low biomass environments are genuine, and not a false positive as a result of cross-contamination [513].

Organic contamination in in situ chemical analysis during spaceflight is always a prime concern, particularly given the low threshold of background organic material, which would amplify any detection of trace organic contaminants. Missions focused on life detection and biosignature preservation that target sensitive parameters, such as amino acid ee, need to prioritize contamination control not only during spacecraft assembly and testing but also instrument operations and data returned for analysis [514]. When determining chirality in organic compounds during life detection missions, contamination could arise from malfunctioning instrumentation [515], sampling [365], or terrestrial contamination [516]. Strategies employed for dealing with contamination include the use of sterilization procedures, e.g., decontamination heaters, UV radiation, chemical, and low-heat plasma ([517] and references therein) and limiting the use of solvents and/or derivatization agents [500].

### 9.4. Future Directions

There are many fields that are interested in and actively researching the chirality of organic and mineral systems, as evidenced by the extensive and growing number of enantioselective techniques (reactions and analytical applications) being developed (Tables 1 and 6). While research in chirality has been dominated by organic synthesis of natural products and bioactive compounds of interest to the pharmaceutical industry, its relevance and importance continues to grow in the field of life detection and origins of life. Research in combining state-of-the-art separation and detection techniques with analog and laboratory studies to constrain conditions in which asymmetric prebiotic reactions are

avored. This would increase our understanding of what life requires, or what conditions should be present for life to be established and to thrive, regardless of whether it is life that is familiar to us or not. In order to address the use of chiral asymmetry as a biosignature, multidisciplinary collaboration is key to facilitating the development of new technology, which could lead to new field-defining instrumentation in the field [124,422,518] and references therein.

## 10. Conclusions

Homochirality is of great interest and importance to prebiotic chemistry as well as those researching the origins of life. Chirality is an important property observed in both organic materials and mineral structures and these structural properties have implications for reactivity and analysis. Understanding both organic and mineral chiral systems is relevant for prebiotic chemistry, as the origins of life had to occur in a geologic context of some kind. We have described asymmetric organic reactions as well as mineral alterations that involve chiral systems to emphasize their relevance to geological processes and highlight the diversity in their applications. Lastly, we recommended some future research directions in the field of chiral organic and mineral systems, in particular, technological and scientific advancements that include: (1) Identifying additional chiral organics and chiral minerals that could be used as biosignatures; (2) focusing on geologically and astrobiologically relevant reactions and systems; and (3) designing instruments that can be modified for spaceflight and remote operation to further the search for life on other worlds.

**Supplementary Materials:** The following supporting information can be downloaded at: <https://www.mdpi.com/article/10.3390/sym14030460/s1>; Table S1: List of chiral minerals identified by their space group.

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