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Key Points:

- We explore the geophysical feasibility of subduction in Europa's ice shell finding that under plausible conditions subduction may occur
- We construct a simple model to explore evolution of porosity and temperature within a slab that is forced to subduct
- Initial porosity and salt content determine if subduction is possible providing constraints on the structure and composition of the ice shell

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Received 14 JUN 2017 Accepted 21 NOV 2017 Accepted article online 4 DEC 2017 Porosity and Salt Content Determine if Subduction Can Occur in Europa's Ice Shell

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Abstract Motivated by recent evidence for subduction in Europa's ice shell, we explore the geophysical feasibility of this process. Here we construct a simple model to track the evolution of porosity and temperature within a slab that is forced to subduct. We also vary the initial salt content in Europa's ice shell and determine the buoyancy of our simulated subducting slab. We find that porosity and salt content play a dominant role in determining whether the slab is nonbuoyant and subduction in Europa's ice shell is actually possible. Generally, we find that initially low porosities and high salt contents within the conductive lid are more conducive to subduction. If salt contents are laterally homogenous, and Europa has a reasonable surface porosity of $\phi_0 = 0.1$, the conductive portion of Europa's shell must have salt contents exceeding ~22% for subduction to occur. However, if salt contents are laterally heterogeneous, with salt contents of ~5%. Thus, we argue that under plausible conditions, subduction in Europa's ice shell is possible. Moreover, assuming that subduction is actively occurring or has occurred in Europa's recent past provides important constraints on the structure and composition of the ice shell.

Plain Language Summary Europa, one of Jupiter's icy moons, has features that look similar to mid-ocean spreading ridges, and recent observations suggest that Europa may also have something similar to subduction zones. This implies that Europa may be the only planetary body other than Earth to have plate tectonics. We construct a computer simulation to track the temperature and compaction of ice as it sinks into warmer underlying ice. When ice warms up, it becomes less dense and easier to compact. The salt content of Europa's plates is also important because salt is denser than ice. The temperature, salt content, and how compacted the ice is determine the density and whether the ice can sink. We find that if the salt content of Europa's ice shell varies by few percent from place to place, then the plates can sink. This supports the idea that something like plate tectonics is occurring on Europa and may tell us about the composition of Europa's ice shell. Our work also implies that the plates will sink all the way to Europa's subsurface ocean. This is important because material from the surface of Europa could act as food for life that may exist in Europa's ocean.

1. Introduction

The surface of Europa, one of Jupiter's icy satellites, is marked by a myriad of extensional tectonic features (e.g., Greenberg et al., 1998; Hoppa et al., 1999). The dearth of craters on Europa implies a young surface age of 40–90 Myr and that Europa is actively resurfaced (Bierhaus, Zahnle, & Chapman, 2009). The ubiquity of extensional tectonics, including dilational bands similar to mid-ocean spreading ridges (Prockter et al., 2002; Sullivan et al., 1998), and the lack of corresponding compressional tectonics have led to the hypothesis that resurfacing on Europa is achieved by subduction or subsumption of portions of the ice shell (Kattenhorn & Prockter, 2014). Detailed plate reconstructions act as strong support for this hypothesis (Kattenhorn & Prockter, 2014). Similar plate reconstructions focused on other regions of Europa also find evidence of subduction (Collins et al., 2016; Perkins, Patterson, & Prockter, 2017).

As a descending plate warms up, it may be integrated into warmer underlying ice, or subsumed. If, however, the plate remains dense and continues to descend, subduction is a more apt term. We will use the term subduction in this paper and later argue why this term may be more appropriate. It is unclear if subduction is initiated and driven by convection (Kattenhorn & Prockter, 2014) or if other driving forces such as nonsynchronous rotation (Geissler et al., 1998; Wahr et al., 2009) or true polar wander (Matsuyama & Nimmo, 2008; Schenk, Matsuyama, & Nimmo, 2008) aid in driving plate motions and subduction. When including plasticity

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Figure 1. Range of (a) possible temperature and (b) normalized porosity within the slab at the moment subduction begins (time = 0). Initial temperature (Figure 1a) and normalized porosity (Figure 1b) plotted as a function of depth for different conductive shell thicknesses, h_{cond} , and assumed viscosities at the base of the conductive portion of the ice shell η_b (solid curves $\eta_b = 10^{15}$ Pa s; dashed curve $\eta_b = 10^{13}$ Pa s).

and yield strengths less than 100 kPa, simulations of convection in Europa's ice shell produce conductive lids and surfaces that are continually overturning or have episodic catastrophic overturns (Showman & Han, 2005). O'Neill and Nimmo (2010) similarly advocate for a relatively low yield strength to explain episodic overturn and high heat flows in the south polar terrain of Saturn's moon Enceladus.

Despite the overturn and recycling of the surface occurring in some Europan convection models (Showman & Han, 2005), we identify major questions about the geophysical feasibility of subduction in Europa's ice shell. The density difference between solid ice and a liquid ocean makes it essentially impossible to subduct a purely conductive ice shell into an underlying ocean. However, a colder conductive ice shell could subduct into a warmer underlying ice layer, since colder ice is denser and would therefore be negatively buoyant relative to surrounding warm ice. Thus, as did Kattenhorn and Prockter (2014), we assume that the conductive portion of the ice shell is subducting into a warm convecting portion of the ice shell (e.g., Nimmo & Manga, 2009) (Figure 1). For a conductive lid with a surface temperature of 100 K, a basal temperature of 260 K, and a thermal conductivity proportional to 1/T, the typical thermal anomaly between the subducting conductive lid and underlying warm convecting ice will be $\Delta T \approx$ 93 K (Nimmo, Pappalardo, & Giese, 2003). The density anomaly associated with this temperature difference is given by $\Delta \rho = -\alpha \rho \Delta T$, where $\alpha = 1.56 \times 10^{-4}$ (T/250 K) K⁻¹ is the temperature-dependent volumetric coefficient of thermal expansion and $\rho = 920 \text{ kg/m}^3$ is a reference density typical for ice (Kirk & Stevenson, 1987). The average density anomaly is thus $\Delta \rho \approx 9$ kg/m³, where the subducting conductive lid is denser than the warmer underlying ice. This simple analysis does not, however, consider porosity in the conductive shell. If the conductive portion of the ice shell is on average more than 1% more porous than the underlying warm convecting ice, the "subducting" shell will be buoyant (the density anomaly associated with porosity differences, $\Delta \phi$, is given by $\Delta \rho = -\rho \Delta \phi$). It is difficult to imagine that subduction could occur if the shell were buoyant.

The porosity at Europa's optical surface may be higher than 95% (Hendrix, Domingue, & King, 2005). The 70 cm radar scattering of Ganymede and Callisto, the other two icy Galilean satellites, imply near surface porosities exceeding 40% (Black, Campbell, & Ostro, 2001). Modeling of 3.5 cm, 13 cm, and 60 cm wave-length radar reflectivities suggests that Europa has a surface layer with appreciable porosity that is meters thick (Black, Campbell, & Nicholson, 2001). Although Europa's near surface has significant porosity, little is known about porosity in Europa's ice shell at depths exceeding a few meters. Recent results of NASA's Gravity Recovery and Interior Laboratory (GRAIL) mission (Zuber et al., 2013) suggest that the Moon's crust is 12% porous to depths of at least a few kilometers (Besserer, Nimmo, & Wieczorek, 2014; Wieczorek et al., 2013). Although porosity on an icy satellite may be quite different than porosity on the Moon, pore space is likely created by fracturing and dilatancy, increase in rock volume with deformation (e.g., Jaeger, Cook, & Zimmerman, 2009), as the ice shell is tectonized. Thus, the

extensional tectonic features covering the surface of Europa (e.g., Greenberg et al., 1998; Hoppa et al., 1999) suggest that the conductive portion of Europa's ice shell may have appreciable porosity. Of course, the porosity and temperature of the "subducting" ice will evolve with time. As the shell subducts, pore space in the ice will eventually collapse and diffusion of heat will erase the thermal anomaly. Because the rate of pore space crushing depends on material viscosity, and thus temperature, the evolution of porosity will be linked to the thermal evolution of the subducting slab. This coupled evolution of porosity and temperature in a subducting ice shell is the focus of this work.

In section 2 we describe our numerical model and how we track the evolution of porosity and temperature of a subducting slab. In section 3.1 we describe our fiducial model and the effect that initial porosity has on the buoyancy of the subducting slab. In section 3.2 we study the role of the temperature structure within Europa's ice shell. In section 3.3 we discuss the effects of bending, subduction dynamics, and uncertainties in the rheology of ice. In section 3.4 we explore the consequence of initial salt content, including lateral variations and the conditions needed to make subduction possible. Finally, in section 4 we discuss implications of our work for the structure and composition of Europa's ice shell.

2. Model Description

We use a simple one-dimensional model to track the evolution of temperature and porosity within a slab as it subducts through a reference ice shell, consisting of a conductive lid and an underlying warm convective layer (Figure 2). Figure 1 shows the range of initial (time = 0) porosity and temperature profiles of the slab used in our models, and Figure 2 presents a schematic of the model. Our simulations track the evolution of an individual parcel or slice of the slab. The parcel is broken into 200 equal-sized cells. At each time step, the depth of each cell is evolved assuming a constant slab velocity and a subduction geometry, as shown in Figure 2. As the depth of the parcel increases, the top temperature boundary condition changes and overburden pressure increases. At each time step, temperature is evolved using Fourier's law of heat conduction and porosity is evolved by tracking removal of pore space by viscous flow. We discuss the model in more detail below.

First, we consider whether a one-dimensional model is a reasonable representation of a subducting slab. Thermal density differences will persist for a timescale τ given by

τ

$$=\frac{h_{\rm cond}^2}{\kappa} \tag{1}$$

(Turcotte & Schubert, 2002), where the thermal diffusivity $\kappa = 10^{-6} \text{ m}^2 \text{ s}^{-1}$ and h_{cond} is the typical length scale of the thermal anomaly (Klinger, 1980; Nimmo et al., 2003). For $h_{cond} = 6$ km, the thickness of the conducting portion of the ice shell in our fiducial model, $\tau \approx 1$ Myr. The distance the subducting slab will travel during this time is given by $l = v_{spread} \tau$, where v_{spread} is velocity of the subducting slab. A comparison of the thickness of the conducting portion of the shell, h_{cond} (i.e., the thickness of the subducting slab in our model), and this distance, *l*, provides an estimate of the importance of thermal diffusion along the direction of slab motion. For example, if $l \approx h_{cond}$, heat diffusion from the top to the bottom of the slab will be comparable to the lateral diffusion of heat along the slab. If, however, $l \gg h_{cond}$, heat will primarily diffuse through the top of the slab and a one-dimensional model is reasonable. To compare *l* and *h*_{cond}, we first need an estimate of the velocity of the subducting slab, v_{spread}. Following Kattenhorn and Prockter (2014), we assume that the slab velocity is equal to the estimated spreading rate at Europa's dilational bands. Based on a mid-ocean spreading ridge model and a relationship between fault block spacing and v_{spread}, Stempel, Barr, and Pappalardo (2005) argue that v_{spread} could be up to 40 mm/yr. Kattenhorn and Prockter (2014) suggest that this rate is consistent with subduction on Europa being a dominant resurfacing mechanism. Convection models with overturning lids predict even higher surface velocities of 100-1000 mm/yr (Showman & Han, 2005). At a spreading rate of 40 mm/yr, v_{spread}τ≈46 km, substantially larger than the thickness of the subducting slab. Thus, a onedimensional conduction model will provide reasonable estimates of temperature evolution within a subducting slab.

We assume an initial surface temperature of $T_s = 100$ K and temperature at the base of the conducting portion of the slab of $T_b = 260$ K. The thermal conductivity of ice is given by Petrenko and Whitworth (1999) as



Figure 2. Geometry of a simple one-dimensional model of a subducting slab. The modeled parcel of material is shown at three times, time = 0, t_{bend} , and t_n . Note that the schematic only shows 5 cells, while our actual model has 200 cells. Additionally, for a time step $\Delta t = 1$ year and spreading rate $v_{spread} = 40$ mm/yr, the cells have a horizontal extent $\Delta x \approx v_{spread} \Delta t = 2.8$ cm while for $h_{cond} = 6$ km the vertical dimension $\Delta y = h_{cond}/n_{cell} = 30$ m, where n_{cell} is the number of cells. The dark blue material at the top left represents the dilational spreading bands. The schematic shows a slab bent through the dip angle $\theta = 45^{\circ}$ along a radius $R_{bend} \approx 2 h_{cond}$ in a time $t_{bend} = R_{bend}\theta/v_{spread}$. After time t_{bend} , slab continues to subduct at a constant dip angle and constant rate, v_{spread} .

$$k = 651 \left(\frac{1K}{T}\right) \,\mathrm{Wm}^{-1}\mathrm{K}^{-1} \tag{2}$$

For a temperature-dependent thermal conductivity, proportional to 1/T, the steady state temperature of the conducting portion of the shell is given by

$$T(z) = T_s \left(\frac{T_b}{T_s}\right)^{\frac{z}{h_{cond}}},$$
(3)

where z is depth and h_{cond} is the assumed thickness of the conductive portion of the shell ($h_{cond} = 6$ km, in our fiducial model) (Nimmo et al., 2003). Below h_{cond} the temperature is assumed to be constant at 260 K (Figure 1) approximating the behavior of nearly isothermal convecting ice (Nimmo et al., 2003). In an actual ice shell the temperature would increase in a thin boundary layer near the interface of the ice shell and ocean (e.g., Hussmann, Spohn, & Wieczerkowski, 2002).

We assume the starting porosity of the ice shell as given in Figure 1. This porosity structure represents an initially constant porosity, ϕ_0 , ice shell that is allowed to evolve for $t_{por} = 65$ Myr, in the middle of estimates of 40–90 Myr for the age of Europa's surface (Bierhaus et al., 2009). Although one could assume that ϕ_0 depends on depth, so little is known about porosity at depth in Europa's ice shell that we argue that the prudent choice is a constant ϕ_0 . This can be thought of as a depth-averaged initial porosity, which is then allowed to evolve through time t_{por} . This reference porosity as a function of depth is described exactly by

$$\phi(z) = \phi_0 \exp\left(-\frac{t_{\text{por}}P}{\eta(z)}\right),\tag{4}$$

where the pressure $P = \rho gz$, $g = 1.31 \text{ m/s}^2$ and $\eta(z)$ is the viscosity as a function of depth. The rheology of ice is often written as an effective Newtonian viscosity (e.g., Nimmo et al., 2003) given by

$$\eta(z) = \eta_b \exp\left(\frac{Q}{R} \left[\frac{1}{T(z)} - \frac{1}{T_b}\right]\right),\tag{5}$$

where the activation energy is Q = 50 kJ/mol (Goldsby & Kohlstedt, 2001), the reference viscosity of the base of the shell is $\eta_b = 10^{15}$ Pa s (for our fiducial model) (Nimmo et al., 2003), *R* is the ideal gas constant, and *T*(*z*) is the temperature as a function of depth (Figure 1). Note that *T* can be substituted for *T*(*z*) in equation (5) to give the temperature-dependent viscosity $\eta(T)$. Our results are rather insensitive to the choice of t_{por} since viscosity $\eta(z)$ is strongly dependent on temperature and depth. Thus, we do not expect a large change from the results presented here even if the subducting material is much older than 65 Myr or if fracturing and

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Figure 3. Evolution of (a) temperature, (b) porosity, and (c) depth in a subducting slab. Material in the slab is colored according to temperature (Figure 3a) and porosity (Figure 3b) as described by the color scale bars. These results are from our fiducial model $h_{cond} = 6$ km, $v_{spread} = 40$ mm/yr, $\eta_b = 10^{15}$ Pa s, and an initial temperature and porosity structure as describe in Figure 1, where $\phi_0 = 0.1$.

dilatancy associated with extensional tectonics (e.g., Greenberg et al., 1998; Hoppa et al., 1999) create porosity on a time scale much shorter than t_{por} .

We refer to the infinitesimal slice of the subducting slab that our model represents as a parcel of the slab. At time zero in our model, the center of the slab moves at a constant rate v_{spread} along a radius of curvature $R_{bend} = 2 h_{cond}$ (in our fiducial model) so that in a time $t_{bend} = R_{bend}\theta/v_{spread}$, the parcel has rotated through the dip angle θ (Figure 2). The depth of each cell is evolved according to this geometry (Figure 3). After t_{bend} , the depth of each of the 200 cells is evolved according to the spreading rate and dip angle of the slab, assumed to be $\theta = 45^{\circ}$ in our fiducial model. The dynamics of a bending plate and dip angle are essentially unconstrained on Europa. In section 3.3, we explore the effect of assumed bending dynamics by varying spreading rate, v_{spread} , R_{bend} , and θ . Although our model is quite simple, it serves as a starting point to explore the geophysical feasibility of subduction in Europa's ice shell.

The temperature boundary condition at the bottom of the slab is assumed to be a constant 260 K. The temperature at the contact of a warmer and cooler object of the same composition will be the average of their respective temperatures. Thus, the top temperature boundary condition of our model,

$$T_{\rm BCtop}(z_{\rm top}) = [T_{\rm ref}(z_{\rm top}) + T_1]/2$$
(6)

where z_{top} is the depth of the top of the parcel, $T_{ref}(z_{top})$ is the temperature in the reference shell (Figure 1) at a depth z_{top} , and T_1 is the current temperature of the topmost cell. In practice, $T_1 \approx T_{ref}$ and the boundary

temperature is very close to T_{ref} . Because the temperature of the reference ice shell T_{ref} is assumed to be constant with time at a given depth, this boundary condition assumes that the cold subducting slab does not alter the temperature of the surrounding material and tends to overestimate heating of the subducting slab. One way to quantify that the effect of assuming T_{ref} is constant with time at a given depth is to assume that the top cell of the parcel does not heat up at all. In this case, the top temperature boundary condition as a function of depth would be given by

$$T_{\text{BCtop-COMP}}(z_{\text{top}}) = [T_{\text{ref}}(z_{\text{top}}) + T_s]/2. \tag{7}$$

If the temperature of the reference ice shell was permitted to evolve, we would expect the top boundary temperature to be intermediate of the results given by equations (6) and (7) (note that equation (7) is only used for comparison; equation (6) is the boundary condition we actually use in our model). The extreme and clearly unphysical assumption for equation (7) would result in at most a 30% difference from the temperature boundary condition we used. Using a one-dimensional model neglects the diffusion of heat along the direction of slab motion. In this way, the one-dimensional model tends to underestimate heating of the slab. This may counteract the overestimate of heating caused by assuming that the temperature of the reference ice shell is unchanging. Thus, our assumed top boundary condition and use of a one-dimensional model are reasonable for our order of magnitude estimates of slab buoyancy.

The evolution of our model through time is achieved through a simple explicit finite difference scheme. At each time step we calculate the heat flow in our model using Fourier's law of heat conduction

$$q = -k \frac{dT(y)}{dy},\tag{8}$$

where *q* is the heat flow per unit area, *k* is the temperature dependent thermal conductivity (equation (2)), *y* is the distance from the top of the parcel (Figure 2), and *T*(*y*) is the temperature as a function of *y* as calculated in the previous time step. The temperature is then evolved using the heat flow calculated from equation (8) and assuming a temperature-dependent specific heat capacity for ice of c = 1925 (*T*/250 K) J/kg/K (Kirk & Stevenson, 1987). With 200 cells and temperature-dependent thermal conductivity, we find that a time step of 1 year produces models that reliably converge. The temperature evolution model was validated against the analytical solution for instantaneous heating of a half-space (Turcotte & Schubert, 2002).

During each time step the porosity in each cell, ϕ , is removed by viscous flow and evolves with time at a rate depending on confining pressure and temperature-dependent viscosity

$$\frac{d\phi}{dt} = -\frac{\phi P}{\eta},\tag{9}$$

given by Fowler (1985), where η is the temperature dependent viscosity described by equation (5), $P = \rho gz$, and z is the current depth of the cell. The porosity ϕ and viscosity η in equation (9) are the porosity and viscosity from the previous time step. The porosity evolution model was validated using the analytical solution for pore space crushing in an initial constant porosity slab (i.e., equations (4) and (5)) (Nimmo et al., 2003).

After the porosity and temperature are updated, the depth of each cell in the parcel is updated according to the geometry shown in Figure 2. Increasing the depth results in an updated overburden pressure $P = \rho gz$ and an updated top temperature boundary condition defined by equation (6) and the current depth of the parcel.

In addition to temperature and porosity, we also consider how salt content may affect the buoyancy of the parcel. To determine if the conducting parcel of the slab is buoyant, we compare the density of the parcel to the density of the reference ice shell at the same depth. We calculate the density difference of each cell using the following equation

$$\Delta \rho = -\alpha \rho_{\rm ice} \Delta T - \rho_{\rm ice} \Delta \phi + \Delta f_{\rm salt} (\rho_{\rm salt} - \rho_{\rm ice}) \tag{10}$$

where ΔT , $\Delta \phi$, and Δf_{salt} are the difference in temperature, porosity, and salt content, respectively, between the cell of interest and material in the reference ice shell at the same depth as the cell of interest (note that ΔT , $\Delta \phi$, and Δf_{salt} are not the changes in temperature, porosity, and salt content with time). The temperaturedependent volumetric coefficient of thermal expansion is given by $\alpha = 1.56 \times 10^{-4}$ (T/250 K) K⁻¹ and $\rho = 920$ kg/m³ is a reference density typical for ice (Kirk & Stevenson, 1987). We assume a salt density $\rho_{salt} = 1444$ kg/m³ appropriate for natron, a potentially dominant component of surface salts based on



Figure 4. Effect of initial porosities, ϕ_0 , on the time evolution of the average density anomaly in the modeled parcel of subducting material. Different curves represent different initial porosities, ϕ_0 , in both the subducting slab and reference shell, as indicated by the legend. All other parameters are unchanged from our fiducial model.

spectral unmixing analyses of McCord et al. (1999). Other spectral unmixing analyses suggest that magnesian sulfates dominate the surface (Shirley et al., 2010; Shirley, Jamieson, & Dalton, 2016). As magnesian sulfates are more dense than natron (up to approximately twice as dense as natron), assuming that natron is the dominant salt component may produce conservative estimates for the density anomaly produced by a given difference in salt content, Δf_{salt} . The average density difference of the parcel, $\Delta \rho_{av}$, is then obtained by summing the density difference of each cell and dividing by the total number of cells. When $\Delta \rho_{av} > 0$ the parcel of the slab is denser than reference material. For reference, if the top of the parcel were instantaneously brought to a depth of h_{cond} (meaning there is no change in temperature or porosity), $\Delta \rho_{av}$ would be ~10 kg/m³, assuming $\phi_0 = 0$, $\Delta f_{salt} = 0$, and $h_{cond} = 6$ km. For the same instantaneous comparison and $\phi_0 = 0.1$, $\Delta \rho_{av} \approx -41$ kg/m³ resulting in a parcel that is buoyant.

3. Results

3.1. Fiducial Model and the Role of Porosity

Figure 3 shows the results of our fiducial model with $h_{\rm cond}$ = 6 km, $v_{\rm spread}$ = 40 mm/yr, and an initial temperature and porosity structure as described in Figure 1, where ϕ_0 = 0.1. In this model, material in the slab

remains colder than the underlying convecting ice into which it is being subducted for ~500 kyr. This timescale is shorter than implied by equation (1), because here we use a temperature-dependent thermal conductivity and heat capacity. The porosity in the slab is completely crushed out by ~260 kyr. At $v_{spread} = 40$ mm/yr, the parcel bends through the dip angle in a time, $t_{bend} \approx 236$ kyr. The slab then descends at a constant angle, and 100 kyr corresponds to increasing in depth by ~2.8 km.

The average density anomaly of the parcel as a function of time (Figure 4) is crucial to understanding how the evolution of temperature and porosity affects the buoyancy of the slab. As we show many plots of average density anomaly versus time, it is prudent to discuss the important aspects of this plot. When the average density anomaly is positive, the modeled parcel of material is nonbuoyant, and we argue that subduction is feasible. When the average density anomaly is negative, the parcel of material is buoyant and subduction is unlikely. In our models, the parcel is forced to subduct with a depth-time history as shown in Figure 1c (when R_{bend} and v_{spread} are unchanged from their values in the fiducial model). Many of our models result in buoyant parcels of material. By exploring the effect of various model parameters, however, we can determine the conditions that produce slabs that are less buoyant and more conducive to subduction. Using the information gained from sensitivity tests performed in other sections and including the possible role of salt content, in section 3.4 we demonstrate that plausible assumptions can produce conditions where subduction is feasible (i.e., nonbuoyant slabs).

It is instructive to first focus on the density anomaly and time evolution of only the thermal anomaly, ΔT (i.e., $\phi_0 = 0$; Figure 4, blue curve). As the parcel of material subducts, the temperature difference between each cell in the parcel and material at the corresponding depth in the reference model increases with time, as does the average density anomaly. The maximum density anomaly is achieved at ~240 kyr. As the parcel continues warming, this density anomaly decreases (Figure 4) until it ultimately disappears at ~500 kyr.

For an initial porosity $\phi_0 = 0.1$, the evolution of porosity dominates the evolution of the average density anomaly with time (Figure 4, black curve). In this case, the porosity difference, $\Delta\phi$, between the subducting slab and the reference material at the same depth increases with time resulting in the parcel of material being buoyant. The initial porosity is rapidly reduced at a depth of 3 km (Figures 1 and 3). Thus, when the modeled parcel of the slab reaches a depth of ~3 km, the difference in porosity between the slab and the reference material reaches a maximum (the top of the slab reaches a depth of 3 km slightly after 190 kyr when $v_{spread} = 40$ mm/yr). After 200 kyr, crushing of pore space begins to dominate and pore space is completely removed by ~260 kyr (Figure 3). After this time, the density anomaly is well described by the thermal



Figure 5. Effect of the thickness of the conductive portion of the shell, h_{cond} , on the time evolution of the average density anomaly in the modeled parcel of subducting material. Different curves represent different conductive layer thicknesses, h_{cond} , in both the subducting slab and reference shell, as indicated by the legend. All other parameters are unchanged from our fiducial model.

evolution and all curves converge (Figure 4). Between $\phi_0 = 0$ and 0.1, the evolution is roughly a linear combination of these two end-members. Motivated by Europa's high surface porosity (Black, Campbell, & Nicholson (2001); Black, Campbell, & Ostro (2001); Hendrix et al., 2005) and the results of NASA's GRAIL mission (Besserer et al., 2014; Wieczorek et al., 2013), we consider $\phi_0 = 0.1$ as our fiducial model. The fiducial model is used as a reference to help understand the effect of changing various model parameters. Models with lower porosity tend to be less buoyant and more favorable for subduction than those with higher porosity. For porosity, less than $\phi_0 = 0.01$, the parcel is nonbuoyant at all times and subduction is feasible (Figure 4).

3.2. Role of Temperature Structure

The thickness of the conducting portion of Europa's ice shell is relatively uncertain (e.g., Nimmo & Manga, 2009). Figure 5 shows how our model results depend on the assumed thickness of the conducting part of Europa's ice shell, h_{cond} . To isolate the thermal effects, we set the bending radius R_{bend} = 12 km and have material at a depth of 3 km (i.e., the bottom of 3 km thick slab; middle of 6 km thick slab, and a third of the way down into the 9 km thick slab) move at a constant rate v_{spread} . This is the same as the fiducial model and ensures that regardless of h_{cond} , the slab bends toward the ultimate dip angle in the same amount of time, t_{bend} . This also

ensures that the evolution of the depth of the top of the parcel and the top temperature boundary condition with time remains the same as we vary h_{cond} . We discuss the role of bending dynamics separately in section 3.3.

Because the surface temperature ($T_s = 100$ K) and basal temperature ($T_B = 260$ K) do not depend on h_{condr} a thicker ice shell corresponds to a lower heat flux and is colder at a given depth than a thinner shell (Figure 1a). The depth at which initial porosity becomes negligible also depends on $h_{\rm cond}$ through equations (4) and (5) (Figure 1b). The mass anomaly created by an increasing difference in porosity with time is nearly the same in these different models. This mass anomaly, however, is averaged over a smaller thickness for models with smaller h_{cond} resulting in an average density anomaly that becomes larger in magnitude at earlier times for smaller h_{cond} . The thickness of the porous layer is nearly proportional to h_{cond} (Figure 1b). Consequently, the minimum average density anomaly occurs later as h_{cond} increases. This is because $\Delta \phi$ is maximized at greater depth for larger $h_{\rm cond}$ since the initial porosity extends to greater depths. We also note that the depths of the minimum average density anomalies are quite insensitive to $h_{\rm cond}$ (Figure 5). These models clearly demonstrate that it takes longer for pore space to be crushed out of the thicker colder subducting slabs, as we might expect. We also see that the thermal anomaly disappears much faster in the thinner, warmer conducting slabs. For $h_{cond} = 3$ km, the thermal anomaly disappears by ~250 kyr (Figure 5), while for $h_{cond} = 6$ km the thermal anomaly persists to ~500 kyr (Figure 2). Although h_{cond} affects the time evolution of slab buoyancy, the lack of a strong effect on the magnitude of buoyancy of the slab seems to suggest the feasibility of subduction as defined in this work is insensitive to the thermal structure of the shell.

3.3. Role of Ice Rheology and Dynamics of Subduction

Following Nimmo et al. (2003), we also consider a range of basal shell viscosities, η_b , from 10^{13} to 10^{15} Pa s to account for uncertainties in the rheology of ice. At the same temperature, pore space crushes out 100 times faster when $\eta_b = 10^{13}$ Pa s as compared to $\eta_b = 10^{15}$ Pa s (equation (9)). Setting $\eta_b = 10^{13}$ Pa s leads to pore space completely crushing out by 220 kyr, whereas it takes 260 kyr when $\eta_b = 10^{15}$ Pa s (Figure 6). Lowering η_b also reduces the depth to which the initial porosity extends (Figure 1), resulting in a lower magnitude for the minimum average density anomaly. These effects together make the slab less buoyant. Thus, models with lower η_b are more conducive to subduction.

The spreading rate, v_{spread} , on Europa is another poorly constrained parameter (Showman & Han, 2005; Stempel et al., 2005). Increasing v_{spread} forces the slab to subduct at a faster rate, meaning the parcel will reach greater depths at earlier times. Thus, a higher v_{spread} results in larger overburden pressures and



Figure 6. Effect of the basal viscosity, η_{b} , on the time evolution of the average density anomaly in the modeled parcel of subducting material. Different curves represent different basal viscosities, η_{b} , in both the subducting slab and reference shell as indicated by the legend. All other parameters are unchanged from our fiducial model.

higher top temperature boundary condition (equation (6)) at any given time. We explore the effect of v_{spread} on the feasibility of subduction by varying v_{spread} from 10 to 100 mm/yr (Figure 7). Note that Figure 7 has time on a logarithmic scale to accommodate the differences in timescale associated with this large range in v_{spread}. Although the initial porosity structure is the same in these models (Figure 1b), we see marked differences in the initial slope of the average density anomaly as a function of time. A steeper slope at higher spreading rates results in a greater mismatch in porosity $\Delta \phi$ at earlier times. For a higher spreading rate, the thermal anomaly disappears faster and pore space is completely crushed out at earlier times compared to models with a lower spreading rate. This faster pore space removal is explained by the top temperature boundary condition, which depends on depth, changing faster with the higher spreading rate. We note that this effect is not captured in a simple application of equation (1) for the lifetime of the thermal anomaly. As $v_{\rm spread}$ decreases from 40 mm/yr to 10 mm/yr, the magnitude of the minimum average density anomaly decreases slightly. Thus, relatively lower spreading rates produce slabs that are less buoyant and slightly more conducive to subduction.

To further test the sensitivity of our models to the dynamics of subduction, we vary the radius of curvature the slab bends through, R_{bend} . As R_{bend}

decreases, cells reach a greater depth at earlier times. Increasing R_{bend} has a similar effect as reducing v_{spread} , resulting in a lower magnitude minimum average density anomaly, which is reached at later times. Both the differences in timescale and the minimum average density anomaly are more sensitive to variation of v_{spread} considered than to changes in R_{bend} (Figures 7 and 8).

Finally, we vary the ultimate dip angle of the slab, θ (Figure 9). The ultimate dip angle of 30, 45, and 60° is reached at a time of 157, 236, and 314 kyr, respectively. The evolution of depth of the parcel is identical for θ = 45 and 60° up until 236 kyr. Because the pore space has been removed from the slab by 236 kyr, the evolution of parcel density is very similar for θ = 45 and 60° (Figure 9). Figure 3c shows that the depth of the bottom of the parcel increases faster once the ultimate dip angle has been reached and bending has ceased. This explains why the evolution of the average density anomaly for θ = 30° looks similar to θ = 45° with a slight shift to earlier times. Ultimately, Figure 9 shows that our results are insensitive to changes in θ .



Figure 7. Effect of the spreading rate, v_{spread} , on the time evolution of the average density anomaly in the modeled parcel of subducting material. All other parameters are unchanged from our fiducial model.

3.4. Role of Salt Content

Next, we consider the role of salts within Europa's ice shell. Radiative transfer modeling of visible and near-infrared spectral data from the NIMS instrument aboard Galileo indicates that salt content on the surface of Europa is heterogeneous and may reach nearly 100% at some locations (McCord et al., 1999). Temperature, grain size, porosity, surface asperities or surface roughness, low albedo components, and trace contaminants, however, can greatly affect a model's salt content estimates (Dalton et al., 2005; Hapke, 1981, 1984; McCord et al., 1999; Milliken & Mustard, 2007a, 2007b). Nimmo et al. (2003) argue that warm ice will be relatively pure because salts will be incorporated in eutectic liquids that drain out of the warm, convecting ice on short timescales. Thus, we might expect that the colder, conductive lid is saltier (and thus denser) than the underlying warm ice. Pits, domes, and small chaos (Michaut & Manga, 2014); ridges (Greenberg et al., 1998); and lobate features (Kattenhorn & Prockter, 2014) have all been suggested to form as the result of cryomagmatism or cryovolcanism. If cryomagmas are sourced from Europa's ocean, which likely has a significant salt content based on the high strength of Europa's



Figure 8. Effect of the radius of curvature through which the slab bends, R_{bend} , on the time evolution of the average density anomaly in the modeled parcel of subducting material. All other parameters are unchanged from our fiducial model.

induced magnetic field (Kivelson et al., 2000), cryomagmatism could enhance the salt content of the conductive lid.

Including a fractional difference in the salt content between the conductive shell and the underlying warm, convecting ice, $\Delta f_{salt} - conv$, results in a lower magnitude minimum average density anomaly (Figure 10). Figure 10 shows that the inclusion of Δf_{salt} – conv is well described by a linear combination of the fiducial model and a constant increase in average density anomaly with depth. The magnitude of this increase in average density anomaly with depth is set by the value of Δf_{salt} – conv. This simple behavior holds until ~290 kyr (Figure 10). At this time, the top of the modeled parcel reaches a depth equal to h_{cond} and the difference in salt content of the parcel and the reference material can no longer increase with depth (i.e., $\Delta f_{\text{salt}} = \Delta f_{\text{salt}}$ - conv for every cell in the slab). For $\phi_0 = 0.1$, $\Delta f_{salt - conv} > 0.22$ is required for the average density anomaly to remain positive at all times meaning the slab would be nonbuoyant (Figure 10). The composition of Europa's ice shell is relatively uncertain, with models of formation and evolution suggesting that the ice shell could be nearly pure water ice or have a significant fraction of

hydrated salts (e.g., Kargel, 2000). Although high, a salt content of 22% is conceivable (e.g., Kargel, 2000), and in such a case subduction would be geophysically feasible.

So far, we have considered porosity, temperature, and salt content variations with depth and assumed that the ice shell is laterally homogeneous. Assuming that the ice shell is laterally homogeneous leads to initial conditions of a parcel with zero average density anomaly. Although variability in salt content at the optical surface seen in the NIMS data does not directly translate to salt content heterogeneity at depth, this observation acts as motivation to consider laterally heterogeneous salt abundance. Compositional heterogeneity may be produced by thermal and compositional diapirism (Pappalardo et al., 1998; Schmidt et al., 2011), intrusive cryomagmatism (Greenberg et al., 1998; Michaut & Manga, 2014), or cryovolcanism (Kattenhorn & Prockter, 2014). Spatially random emplacement of cryomagma would tend to increase the salt content and density of Europa's ice shell over time, making older regions of the ice shell denser.

To explore the effect of lateral compositional heterogeneity, we include a fractional difference between the salt content of the subducting slab and the conductive portion of the average reference ice shell, $\Delta f_{salt - ref}$. A $\Delta f_{salt - ref} = 0.01$, for example, means that the subducting slab is 1% saltier than the average



Figure 9. Effect of ultimate dip angle of the slab, θ , on the time evolution of the average density anomaly in the modeled parcel of subducting material. All other parameters are unchanged from our fiducial model.

ice shell and initially ~5 kg/m³ denser than the surrounding material (equation (10)). This means that the initial state of our model is one where the parcel is nonbuoyant. When material in the slab is deeper than h_{cond} , the density anomaly from salt is $\Delta f_{salt} = \Delta f_{salt - conv}$, the same as in models shown in Figure 10, which do not include lateral heterogeneity. For material above h_{cond} , however, $\Delta f_{salt} = \Delta f_{salt - ref}$ (for the models in Figure 10 this material has $\Delta f_{salt} = 0$). Thus, when $\Delta f_{salt - conv}$ is kept the same, changes to $\Delta f_{salt - ref}$ only change the average density anomalies out to a time of ~290 kyr. Since all porosity is crushed out by 290 kyr, a comparison of the red curve from Figure 11a and the purple curve from Figure 10 (both corresponding to $\Delta f_{salt - conv} = 0.05$) demonstrate that models with different $\Delta f_{salt - ref}$ have the same average density anomaly after 290 kyr.

Because $\Delta f_{\text{salt} - \text{ref}} > 0$ results in a positive average density anomaly at time zero and is initially nonbuoyant, we define a time-averaged density anomaly to assess whether subduction is feasible. The time-averaged density anomaly at a given time, $t_n = n\Delta t$, is given by



Figure 10. Effect of salt content, $\Delta f_{salt} - conv$, on the time evolution of the average density anomaly in the modeled parcel of subducting material. Assumed fractional differences in the salt content of the conducting and convecting portion of the ice shell, $\Delta f_{salt} - conv$, in both the subducting slab and reference shell are indicated by the legend (e.g., $\Delta f_{salt} - conv = 0.1$ means that there is 10% more salt in the subducting slab than the underlying convecting ice). All other parameters are unchanged from our fiducial model.

$$\Delta \rho_{\text{tav}}(t_n) = \sum_{i=1}^n \Delta \rho_{\text{av}}(t_i)/n \tag{11}$$

where $\Delta \rho_{av}(t_i)$ is the spatially averaged density anomaly used throughout this paper at a given time step $t_i = i\Delta t$ and Δt is our time step of 1 year. We define models where subduction is feasible as those with a time-averaged density anomaly that never becomes negative at any time t_n . Because the position and depth of the slab evolve with time (Figures 2 and 3c) this time average is the same as a spatial average over the slab from the location where subduction initiates (t = 0) to the position and depth at time t_n . This ensures that from the initiation of subduction, t = 0, to a time of interest t_n , even if some portions of the slab are buoyant, the slab as a whole is always over-dense and tends to sink. Including lateral heterogeneous salt content allows for slabs that are nonbuoyant. For example, a slab with an initial porosity $\phi_0 = 0.05$, salt content of 2% ($\Delta f_{salt - conv} = 0.02$), and 1% more salt than the reference conductive lid ($\Delta f_{salt - ref} = 0.01$) will be nonbuoyant. Lower η_b and ϕ_0 result in higher time-averaged density anomalies making subduction more likely. Higher $\Delta f_{salt - conv}$ and $\Delta f_{salt - ref}$ also result in higher time-averaged density anomalies. Figure 11 shows a range of parameters for which the time-averaged density anomaly remains positive at all times and subduction is possible. For a plausible combination of parameters like $\phi_0 = 0.1$, η_b = 10^{13} Pa s, $\Delta f_{salt\ -\ conv}$ = 0.05, and $\Delta f_{salt\ -\ ref}$ = 0.025 the time-

averaged density anomaly remains positive and subduction is possible.

In Figure 11 we show models where the modeled parcel of material is buoyant at some times but the slab is always nonbuoyant. It is also possible to construct simulations with plausible parameters in which our modeled parcel of material never becomes buoyant (Figure 12). For example, when $\phi_0 = 0.05$, $\eta_b = 10^{13}$ Pa s, $\Delta f_{\text{salt} - \text{conv}} = 0.04$, and $\Delta f_{\text{salt} - \text{ref}} = 0.02$, the parcel is always nonbuoyant and subduction is feasible regardless of arguments about the buoyancy of the slab as whole (i.e., time averages used above). For larger salt contents and lower porosities than those shown in Figure 12, the modeled parcel will be even more dense and more likely to sink.



Figure 11. Time evolution of (a) the average density anomaly in the modeled parcel of subducting material and (b) time-averaged density anomaly of the subducting slab. Combinations of difference in salt content between the conductive slab and warm convecting ice, $\Delta f_{\text{salt} - \text{conv}}$, difference in salt content between the conductive slab and warm convecting ice, $\Delta f_{\text{salt} - \text{conv}}$, difference in salt content between the conductive slab and warm convecting ice, $\Delta f_{\text{salt} - \text{conv}}$, difference in salt content between the conductive slab and the conductive part of the reference shell, $\Delta f_{\text{salt} - \text{ref}}$, initial porosity, ϕ_0 , and viscosity at the base of the conductive shell, η_b , are plotted as indicated by the legend. All other parameters are unchanged from our fiducial model.



Figure 12. Time evolution of the average density anomaly in the modeled parcel of subducting material showing conditions where the parcel is always nonbuoyant. Combinations of difference in salt content between the conductive slab and warm convecting ice, $\Delta f_{salt - conv}$, difference in salt content between the conductive slab and the conductive part of the reference shell, $\Delta f_{salt - ref}$, initial porosity, ϕ_0 , and viscosity at the base of the conductive shell, η_b , are plotted as indicated by the legend. All other parameters are unchanged from our fiducial model.

4. Discussion

In cases where the slab is nonbuoyant, we suggest that the plate will be subducted rather than subsumed (i.e., incorporated into surrounding material as the thermal anomaly diffuses). Even after a thermal anomaly is erased, most models in which subduction is feasible (Figures 11 and 12) have a density anomaly associated with salt content. Unless salts can drain away from the slab on a time scale of ~1 Myr, this material will continue to descend. Moreover, convection models with overturning lids have high surface velocities exceeding 100 mm/yr (Showman & Han, 2005). In these models, lid material descends to the base of the ice shell before the thermal anomaly can diffuse and the scenario is more consistent with subduction than subsumption.

Our model greatly oversimplifies the process of subduction. Ideally, we would model the evolution of porosity, temperature, and salt content in a dynamic viscoelastic model or in a convection model that includes plastic yielding (Showman & Han, 2005). In such a model, the evolution of the slab would affect the slab dynamics, including slab bending and evolution of the dip angle. In our model, however, we enforce a constant rate of bending and subduction even if the slab is buoyant. A dynamic viscoelastic model could also include a more sophisticated strain rate and temperature-dependent rheology for ice (Durham, Kirby, & Stern, 1997; Durham & Stern, 2001; Goldsby & Kohlstedt, 2001). This type of model could also track the evolution

of temperature without enforcing strict, but reasonable, temperature boundary conditions at the top and bottom of the slab. These proposed simulations would be computationally expensive, and exploring a large parameter space would be more difficult than with our simple one-dimensional model. Considering that the evolution of porosity is strongly dependent on temperature, but only weakly dependent on overburden pressure (equations (4) and (5)), we do not expect the dynamics of plate bending and subduction to have a strong effect on the buoyancy of the subducting slab. Regardless of these simplifications, our modeling shows that porosity and salt content ultimately determine whether subduction in the ice shell is feasible. We argue that porosity and salt content should be considered in any model of ice shell overturn and subduction on Europa or elsewhere (e.g., O'Neill & Nimmo, 2010; Showman & Han, 2005).

Another important issue is that the initial porosity and composition in Europa's ice shell are largely unconstrained. Our modeling shows that relatively small lateral variations in salt content ($\Delta f_{salt - ref} = 0.025$) make subduction in Europa's ice shell possible even when porosity is relatively high ($\phi_0 = 0.1$) and differences in salt content between the conducting portion of the ice shell and underlying warm ice are reasonable ($\Delta f_{salt - conv} = 0.05$). Increasing lateral variation in salt content, overall salt content of the conductive lid, and decreasing porosity all increase the average density anomaly of the slab and make subduction more likely. We note that although we did not explore the effects directly, lateral variations in temperature and porosity would have a similar effect to lateral variations in salt content, producing conditions more favorable to subduction.

Possible mechanisms for driving plate motions and initiating subduction are convection (Showman & Han, 2005) and nonsynchronous rotation (Geissler et al., 1998). Our model is independent of these mechanisms and if we assume that subduction is occurring, or has occurred in the recent past, our results provide useful constraints on the composition and structure of Europa's ice shell. Subduction in the ice shell can account for the abundance of extensional features, a lack of compressional features, and young age of Europa's surface (Kattenhorn & Prockter, 2014). Demonstrating that subduction is geophysically feasible under certain conditions allows for further consideration of this exciting possibility: something like plate tectonics on a body other than Earth (subduction likely occurs on Venus, but it lacks lithospheric plates and mid-ocean spreading ridge analogs (Davaille, Smrekar, & Tomlinson, 2017; Sandwell & Schubert, 1992)). Exchange of surface materials with the subsurface ocean helps determine Europa's astrobiological potential (Chyba & Phillips, 2001; Gaidos, Nealson, & Kirschvink, 1999). Subduction in Europa's ice shell would enhance the rate at which

surface material can be delivered to its subsurface ocean, possibly producing conditions that are more conducive to life (Kattenhorn & Prockter, 2014). A better understanding of how salts and porosity are distributed in the ice shell may also shed light on other important processes at work on Europa including diapirism and cryomagmatism (Michaut & Manga, 2014; Pappalardo et al., 1998; Schmidt et al., 2011).

5. Conclusions

We have explored the geophysical feasibility of subduction in Europa's ice shell using a one-dimensional finite difference model that tracks the evolution of temperature and porosity within a subducting slab. We tested the effects of porosity, salt content, conductive lid thickness, basal viscosity, spreading rate, bending radius, and dip angle on the buoyancy of a subducting slab. We found that the buoyancy of a subducting ice slab is most sensitive to salt content and initial porosity. Higher salt contents, lateral variations in salt content, and lower porosity result in slabs that are less buoyant. We find that slabs with salt contents of 5% ($\Delta f_{salt} - conv = 0.05$) and 2.5% more salt than the reference conductive lids ($\Delta f_{salt} - ref = 0.025$) are nonbuoyant even when they have relatively high initial porosities of $\phi_0 = 0.1$. Our work shows that subduction is geophysically feasible under plausible conditions and supports Kattenhorn and Prockter's (2014) geology-based hypothesis of subduction in Europa's ice shell. Because compositional density variations can remain even after the slab has reached the same temperature as warm ice, material will continue to descend and may transport surface materials to Europa's subsurface ocean.

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